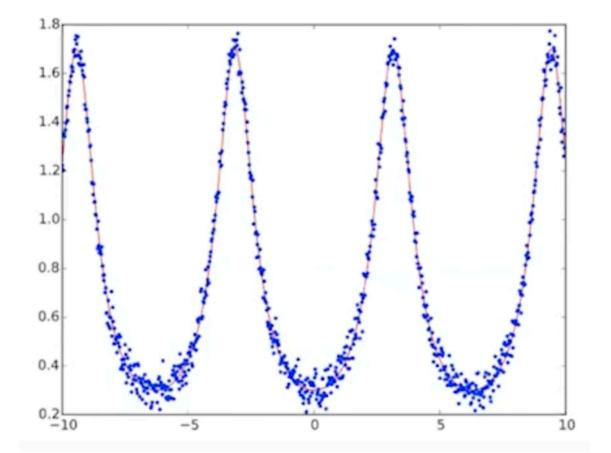
Symbolic Regression and Polynomial Optimization in Scientific Discovery

Karan Srivastava

Aim

• To discover meaningful laws of nature from experimental data

i.e. Given some data $\{(x_1, \ldots, x_n, y)_i\}_{i \in \mathscr{D}}$ we want to find a function f so that $y = f(x_1, \ldots, x_n)$ for each data point.



$$r = \frac{a(1 - e^2)}{1 + e\cos(\theta_1 - \theta_2)}$$

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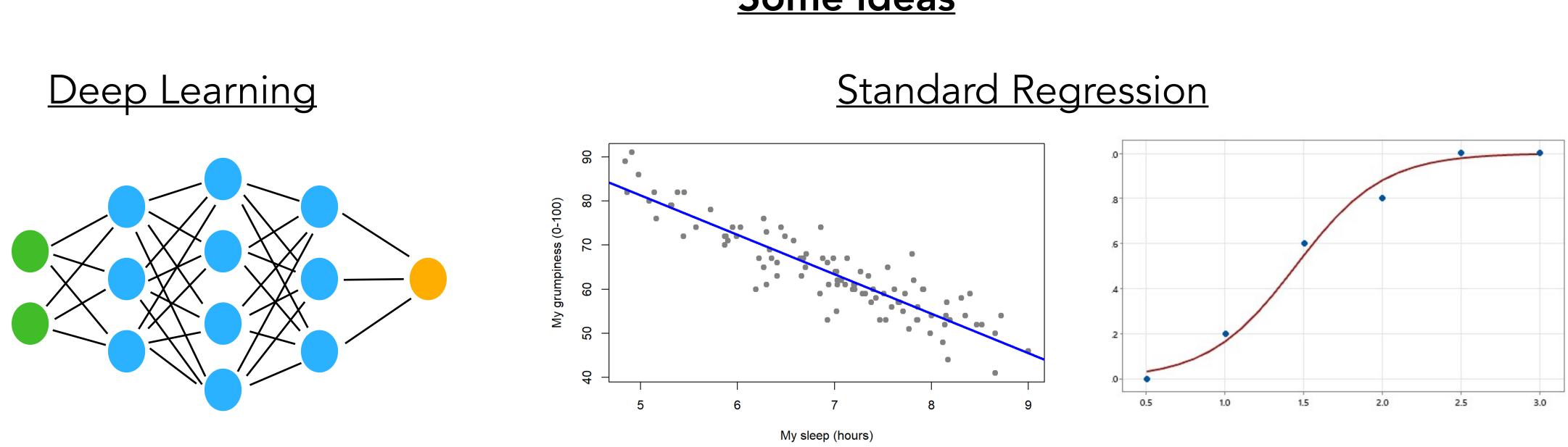
Deep Learning

Pros: Good for discovering patterns in data Drawback: Model itself is uninterpretable

<u>Some ideas</u>

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<u>Some ideas</u>

Pros: Models are very interpretable Drawback: Functional form is fixed

Approach 1: Symbolic Regression

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In particular, we will look at the approach in this paper:

AI Feynman: a Physics-Inspired Method for Symbolic Regression

Silviu-Marian Udrescu, Max Tegmark^{*}

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(Dated: Published in *Science Advances*, 6:eaay2631, April 15, 2020)

A core challenge for both physics and artificial intelligence (AI) is *symbolic regression*: finding a symbolic expression that matches data from an unknown function. Although this problem is likely to be NP-hard in principle, functions of practical interest often exhibit symmetries, separability, compositionality and other simplifying properties. In this spirit, we develop a recursive multidimensional symbolic regression algorithm that combines neural network fitting with a suite of physics-inspired techniques. We apply it to 100 equations from the Feynman Lectures on Physics, and it discovers all of them, while previous publicly available software cracks only 71; for a more difficult physics-based test set, we improve the state of the art success rate from 15% to 90%.

I. INTRODUCTION

In 1601, Johannes Kepler got access to the world's best data tables on planetary orbits, and after 4 years and about 40 failed attempts to fit the Mars data to various ovoid shapes, he launched a scientific revolution by discovering that Mars' orbit was an ellipse [1]. This was an example of symbolic regression: discovering a symbolic expression that accurately matches a given data set. More specifically, we are given a table of numbers, whose rows are of the form $\{x_1, ..., x_n, y\}$ where $y = f(x_1, ..., x_n)$, and our task is to discover the correct symbolic expression for the unknown mystery function f, optionally including the complication of noise.

Growing data sets have motivated attempts to automate such regression tasks, with significant success. For the

search space characterizes many famous classes of problems, from codebreaking and Rubik's cube to the natural selection problem of finding those genetic codes that produce the most evolutionarily fit organisms. This has motivated genetic algorithms [2, 3] for targeted searches in exponentially large spaces, which replace the abovementioned brute-force search by biology-inspired strategies of mutation, selection, inheritance and recombination; crudely speaking, the role of genes is played by useful symbol strings that may form part of the sought-after formula or program. Such algorithms have been successfully applied to areas ranging from design of antennas [4, 5] and vehicles [6] to wireless routing [7], vehicle routing [8], robot navigation [9], code breaking [10], discovering partial differential equations [11], investment strategy [12], marketing [13], classification [14], Rubik's cube [15], program synthesis [16] and metabolic networks [17].

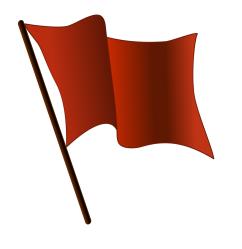
• Given some data $\{(x_1, \ldots, x_n, y)_i\}_{i \in \mathcal{D}}$ we want to find a function f so that $y = f(x_1, \ldots, x_n)$ for each data point when the functional form of f is unknown.

• In particular, given a set S of symbols (e.g. $+, -, \div, \sqrt{,}$ etc), find a function (string) f built from these symbols so that $y = f(x_1, \ldots, x_n)$ fits the data.

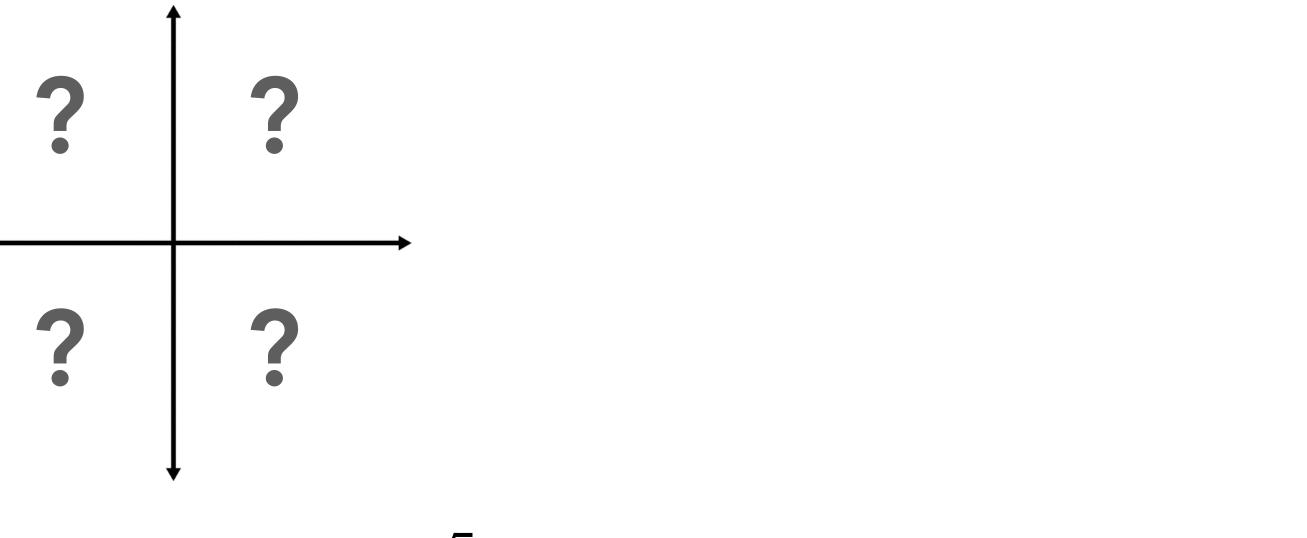
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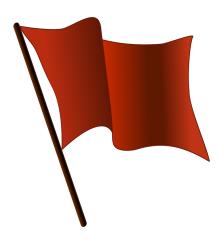


For generic functions $f(x_1, \ldots, x_n)$, this is NP hard





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- Search grows exponentially with the number of symbols.
- Brute force search becomes infeasible very quickly

$$f = \frac{e^{-\frac{x^2}{2}}}{\sqrt{2\pi}} \sim 30 \text{ years,} \quad \omega = \frac{1 + \frac{v}{c}}{\sqrt{1 - \frac{v^2}{c^2}}} \sim$$

For generic functions $f(x_1, \ldots, x_n)$, this is NP hard

 10^6 years, and so on.

| Feynman | Equation |
|---------|---|
| eq. | |
| I.6.20a | $f = e^{-\theta^2/2} / \sqrt{2\pi}$ |
| I.6.20 | $\int f = e^{-\frac{\theta^2}{2\sigma^2}} / \sqrt{2\pi\sigma^2}$ |
| I.6.20b | $\int f = e^{-\frac{(\theta - \theta_1)^2}{2\sigma^2}} / \sqrt{2\pi\sigma^2}$ |
| I.8.14 | $d = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2}$ |
| I.9.18 | $F = \frac{Gm_1m_2}{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2}$ |
| I.10.7 | $m = \frac{\binom{w_2 - w_1}{m_0}}{\sqrt{1 - \frac{v^2}{c^2}}}$ |
| I.11.19 | $A = x_1 y_1 + x_2 y_2 + x_3 y_3$ |
| I.12.1 | $F = \mu N_n$ |
| I.12.2 | $F = \frac{q_1 q_2}{4\pi \epsilon m^2}$ |
| I.12.4 | $\begin{bmatrix} - & 4\pi\epsilon r^2 \\ E_f = \frac{4\pi\epsilon r^2}{4\pi\epsilon r^2} \end{bmatrix}$ |
| I.12.5 | $F = q_2 \tilde{E}_f$ |
| I.12.11 | $F = q(E_f + Bv\sin\theta)$ |
| I.13.4 | $K = \frac{1}{2}m(v^2 + u^2 + w^2)$ |
| I.13.12 | $U = \tilde{G}m_1m_2(\frac{1}{r_2} - \frac{1}{r_1})$ |
| I.14.3 | U = mgz |
| I.14.4 | $U = \frac{k_{spring}x^2}{2}$ |
| I.15.3x | $x_1 = \frac{x - ut}{\sqrt{1 - u^2/c^2}}$ |

BUT SCIENCE IS NOT GENERIC

following ways:

- 1. Units: f and its variables have to be dimensionally consistent
- 2. Low degree polynomials: Parts of f tend to have low degree polynomials
- 3. Compositionality: f is a composition of elementary functions
- 4. Smoothness: f is continuous and often analytic on its domain
- 5. Symmetry: f comes with translational, rotational, and scaling symmetry with respect to some variables

6. Separability: f can be written as the sum or product of two parts with no common variables

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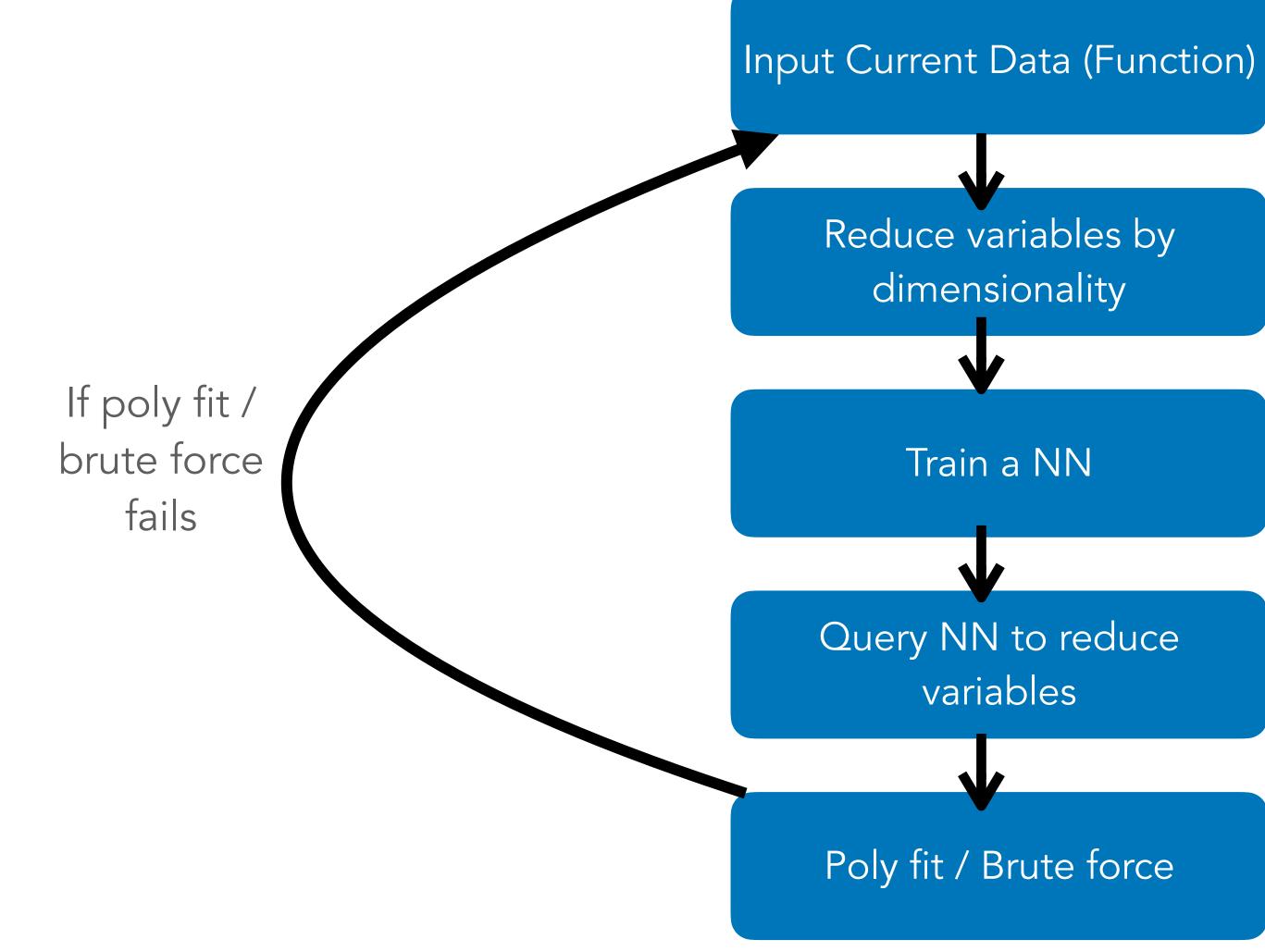
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- 3. Compositionality: Can use brute force for parts of the function
- 4. Smoothness: *f* can be learned by a NN
- 5. Symmetry: Can be tested by a neural network trained on the data
- 6. Separability: Can be tested by a neural network trained on the data

The idea:



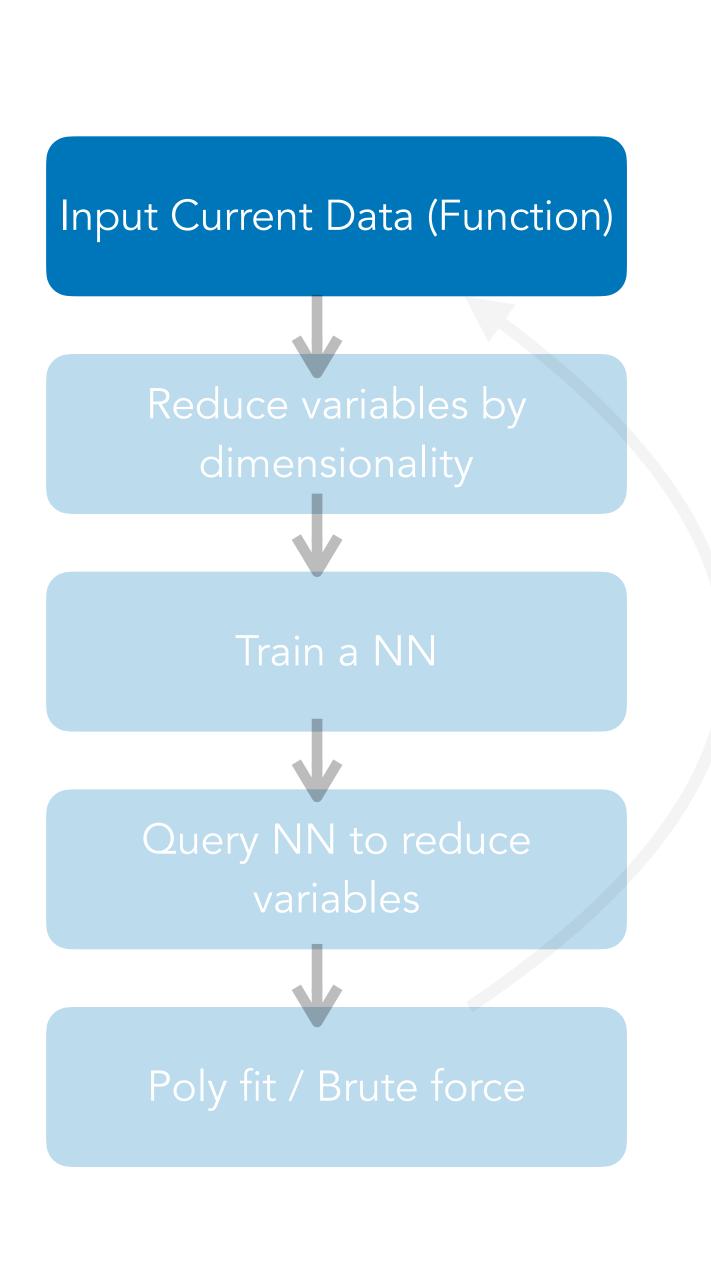
If poly fit / brute force succeeds



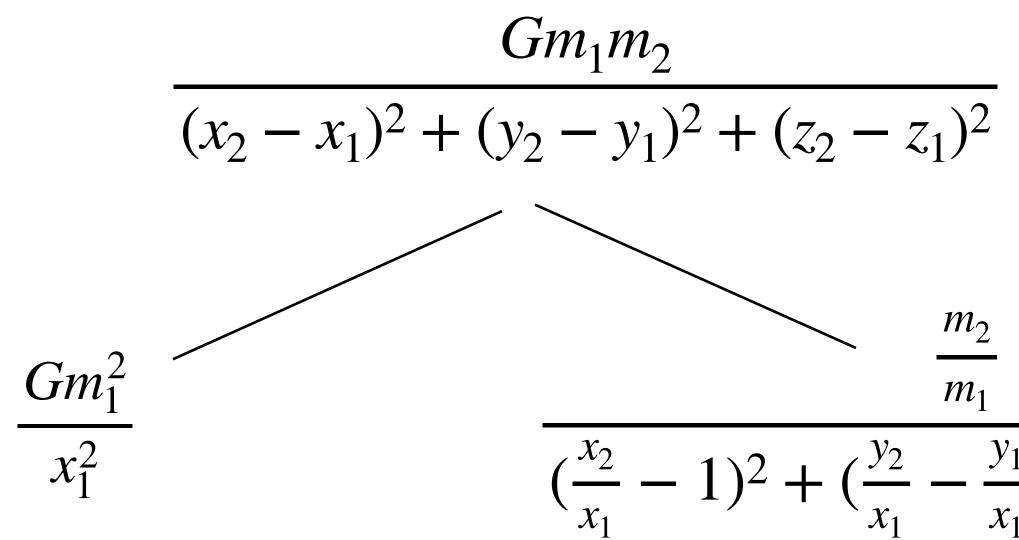
 Gm_1m_2

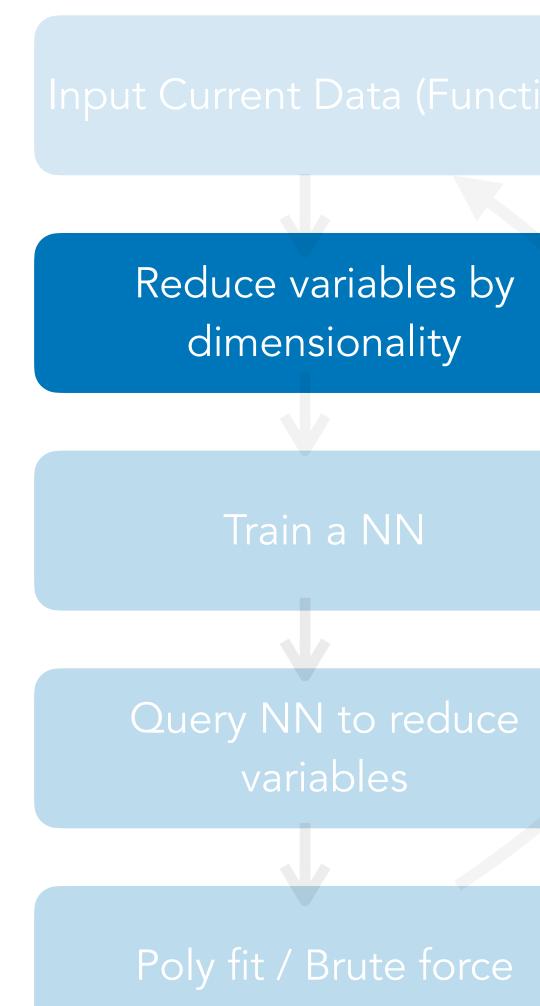
 $(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2$

| 1.1431209959193709 | 2.7700644791483753 | 1.7508575540193472 | 0.234528950 |
|--------------------|--------------------|--------------------|-------------|
| 2.4680655653881054 | 2.2073166947348444 | 1.7761705838854234 | 0.159194033 |
| 2.7621479700455853 | 1.4168131204210188 | 1.5378176974809339 | 0.144293373 |
| 1.9536888384746354 | 2.7336267945043491 | 1.2592849110534683 | 0.153600145 |
| 2.1532278876457527 | 1.5016008010765851 | 1.4218686278023172 | 0.185149407 |
| 1.9899434091665062 | 1.4250958039594244 | 2.5409132056932424 | 0.171314745 |
| 1.2841783534277345 | 2.5038413591290976 | 1.2255232096430668 | 0.189284395 |
| 2.3550853261290494 | 2.2555822345853405 | 1.4525468706453792 | 0.159829295 |
| 2.7529820467784543 | 2.6405850369222492 | 1.6148891450043024 | 0.135195987 |
| 1.2043936184306594 | 1.4441117081403013 | 1.4546229392136278 | 0.331226503 |
| 1.3423962980280737 | 2.0552387587225684 | 2.8071816262301414 | 0.254036157 |
| 1.4156715177406782 | 1.0577553334831364 | 2.5122383795854709 | 0.166238136 |
| 2.0187117984150182 | 1.0672568295716016 | 1.4729243386484654 | 0.193671994 |
| 1.5109507435415031 | 2.5932595628763617 | 2.9492633493443927 | 0.256805829 |
| 2.5078719313855347 | 1.0771974670079432 | 2.7296063077362385 | 0.128036442 |
| 1.4846972064278652 | 2.7401242687034313 | 1.3356870260708698 | 0.171777990 |
| 2.1398682204221178 | 1.5010412666279194 | 2.0784556152016664 | 0.179767914 |
| 1.7000728811732311 | 1.5155671571998763 | 1.5295642030291683 | 0.234653916 |
| 2.3516545980830331 | 1.0329627134609363 | 2.4473327905843174 | 0.141575743 |
| 2.6157973581192184 | 2.0432144044920815 | 1.4576627239471807 | 0.148738971 |
| | | | |

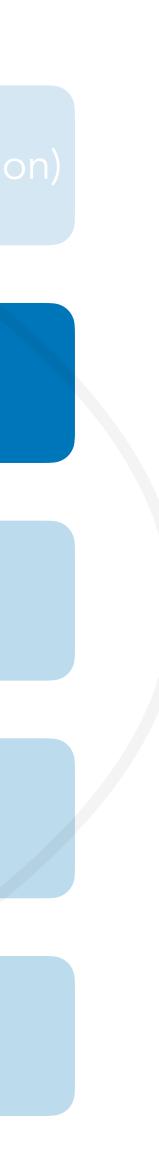


Idea: Because f is dimensionally consistent, we can factor it into a component that encodes the dimension and a dimensionless term



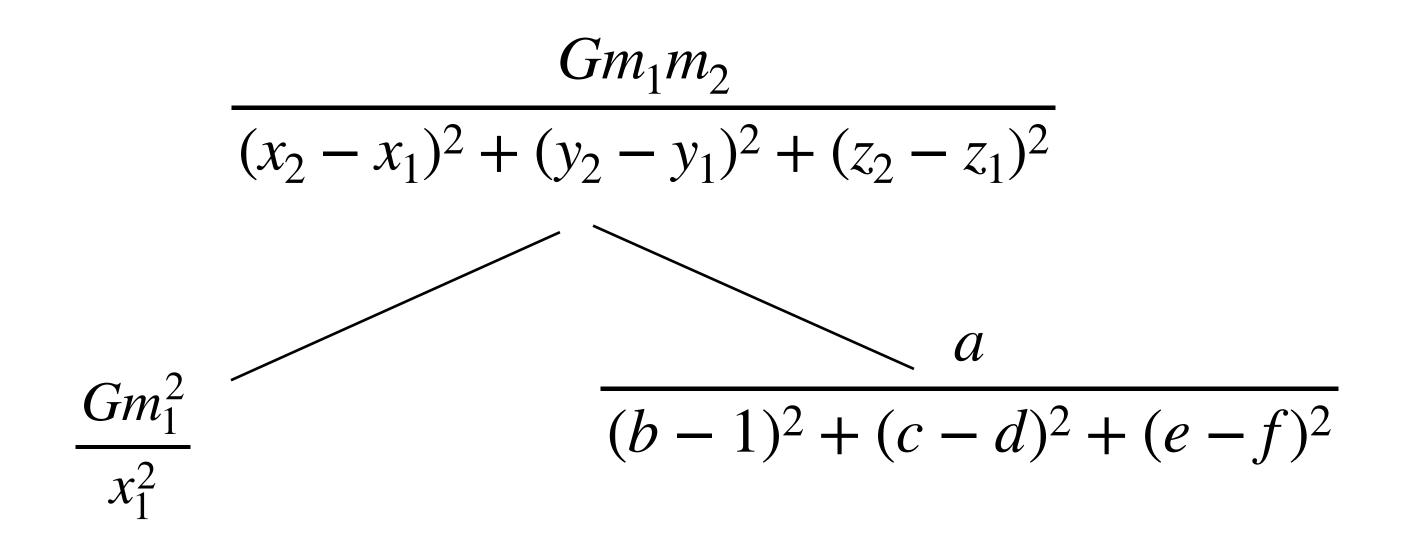


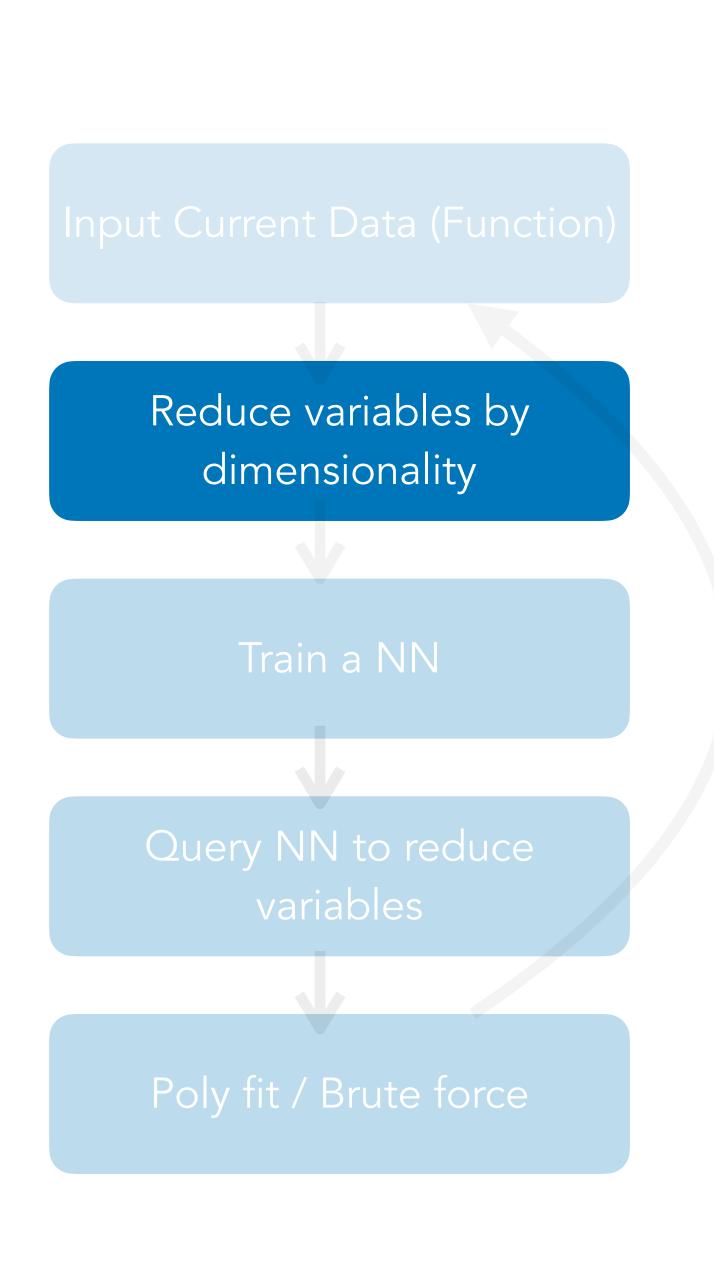
$$(\frac{z_1}{z_1})^2 + (\frac{z_2}{x_1} - \frac{z_1}{x_1})^2$$



Idea: Because f is dimensionally consistent, we can factor it into a component that encodes the dimension and a dimensionless term

This reduces the number of variables to consider.



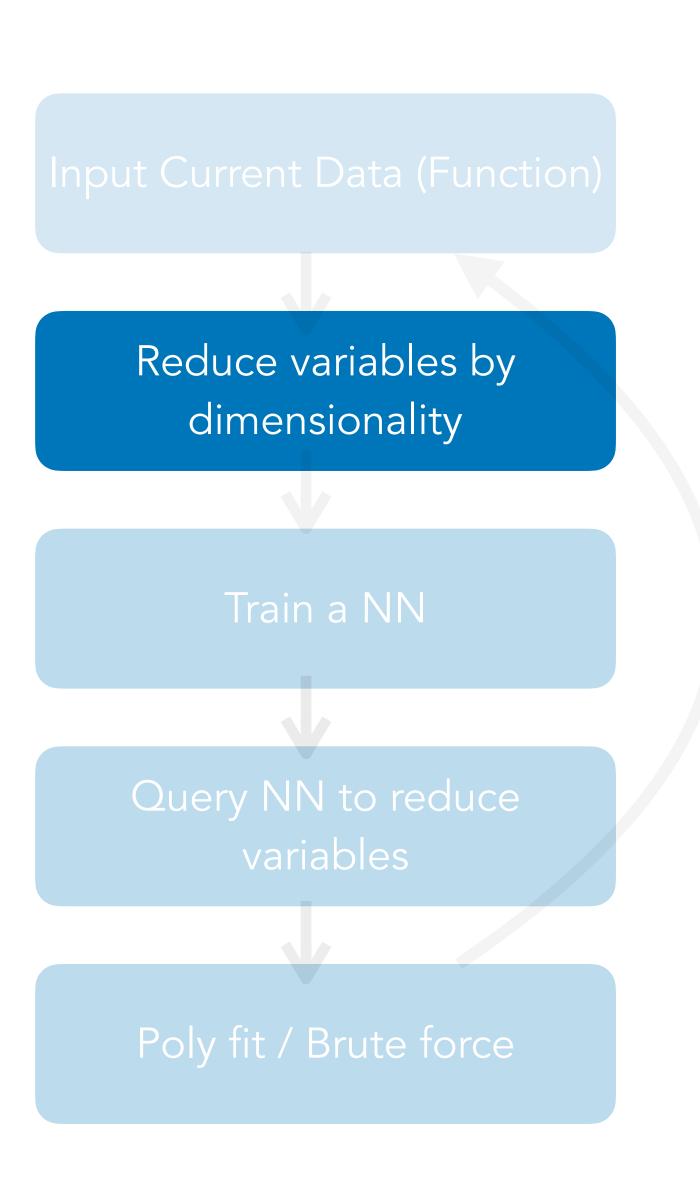


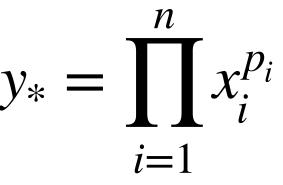
Idea: Because f is dimensionally consistent, we can factor it into a component that encodes the dimension and a dimensionless term

Represent each variable x_i by a 5 integer vector $\mathbf{u_i}$ corresponding to the fundamental units of the vector (being meter, second, kilogram, kelvin, volt). Let \mathbf{M} be a matrix whose *i*th column is $\mathbf{u_i}$. Define \mathbf{b} to be the corresponding vector for y.

Let p be a solution to $\mathbf{M}p=b$ and \mathbf{U} be a basis for the null space of \mathbf{M} Then apply

$$x_i \mapsto \prod_{i=j}^n x_j^{U_i j}, y \mapsto \frac{y}{y_*}, \text{ where } y$$





Train a NN on the transformed columns. This is a black box oracle that we can query.

 $\frac{a}{(b-1)^2 + (c-d)^2 + (e-f)^2}$

Architecture used:

Dimension of layers: (128,128,128,64,64,64) Epochs: 100 Learning rate: 0.005 Batch Size: 2048 Rms loss and Adam optimization Weight decay: 10⁻² Train a NN



Test for symmetry, separability, and other properties and reduce variables accordingly

$$a$$

$$(b-1)^{2} + (c-d)^{2} + (e-f)^{2}$$

$$| \text{ Translational Symm}$$

$$a$$

$$(b-1)^{2} + g^{2} + (e-f)^{2}$$

$$| \text{ Translational Symm}$$

$$a$$

$$(b-1)^{2} + g^{2} + h^{2}$$

$$Multiplicative Separability$$

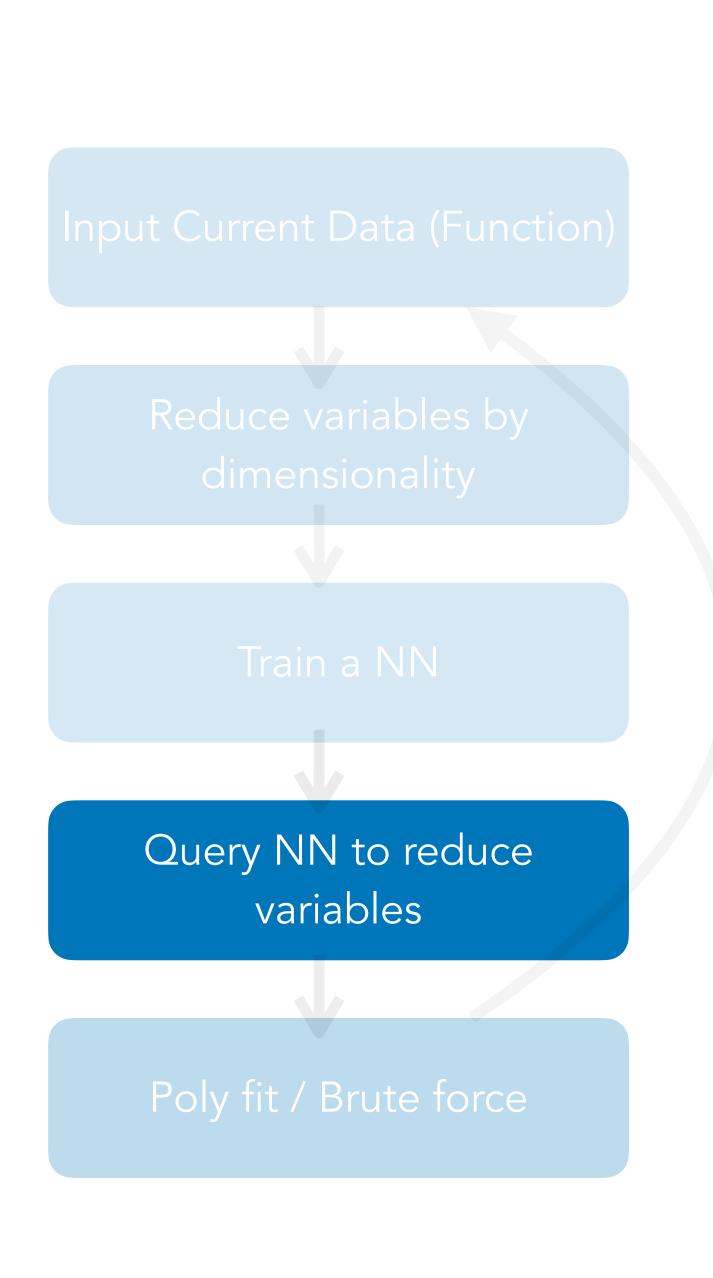
$$a$$

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 $(b-1)^2 + g^2 + h^2$



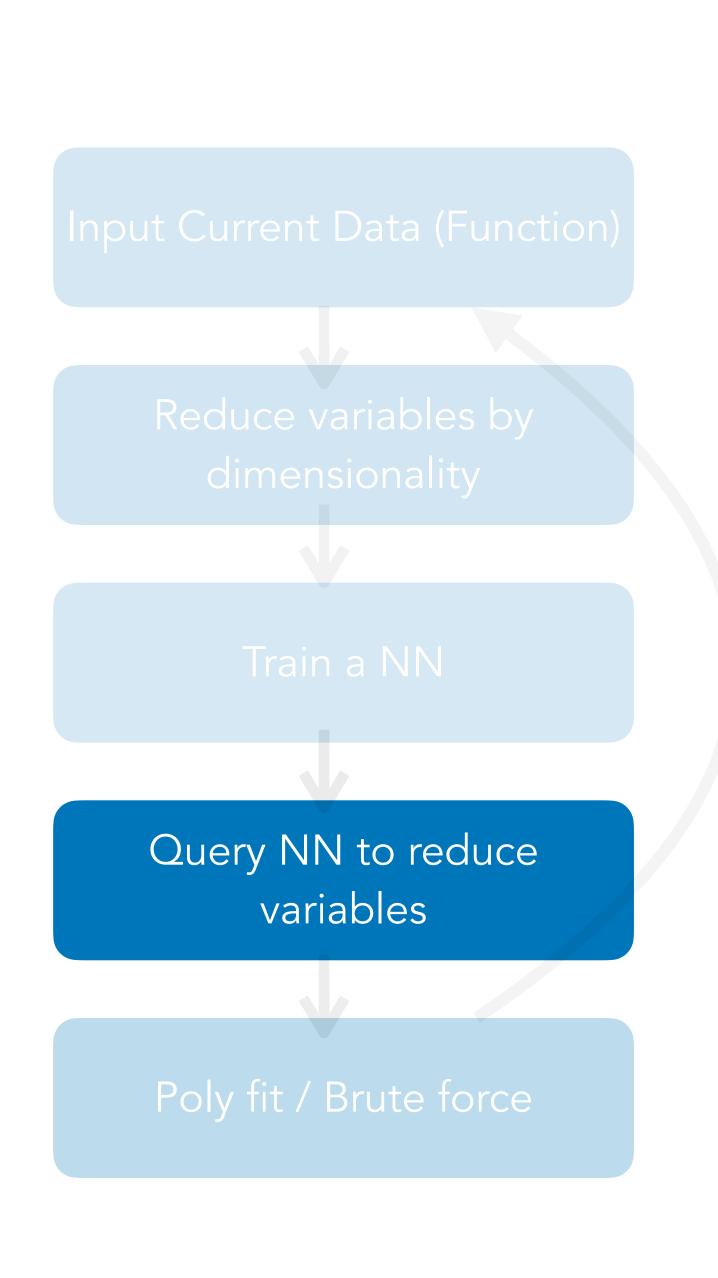
Symbolic Regression - Example **Translational Symmetry**

Let $F(x_1, \ldots, x_n)$ be the function learned by the neural network. Then translational symmetry with respect to x_1 and x_2 corresponds to

$$F(x_1 + a, x_2 + a, x_3, \dots, x_n) - F(x_1, x_n)$$

If the difference is $< \epsilon_{sym}$, then apply $x_1 \mapsto x_2 - x_1$

 $x_2,\ldots,x_n)\approx 0$



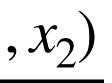
Multiplicative Separability

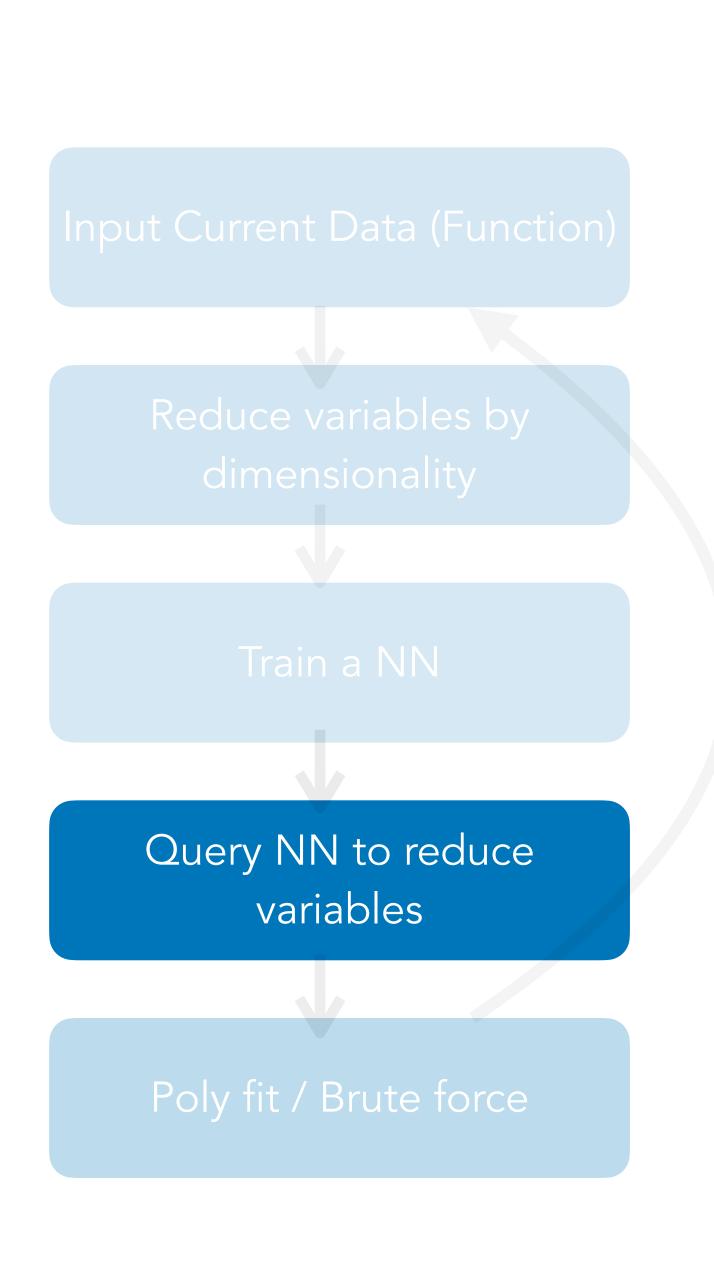
f is multiplicatively separable in x_1, x_2 if f can be factored as $f(x_1, x_2) = g(x_1)h(x_2)$ With no common variables.

To test this, pick constants c_1, c_2 and check if

$$f(x_1, x_2) = \frac{f(x_1, c_2) \cdot f(c_1, c_2)}{f(c_1, c_2)}$$

Up to ϵ_{sep} . If so, then separate the problem of learning f into two subproblems of learning g and h





Translational Symmetry

Multiplicative Separability

Rotational Symmetry

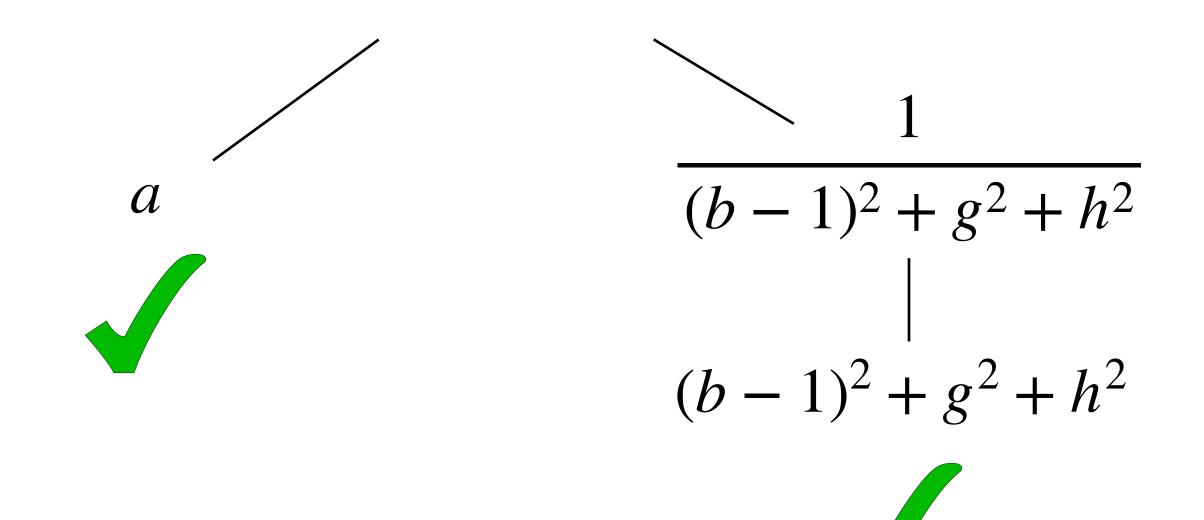
Scaling Symmetry

Additive Separability

Query NN to reduce variables



Test brute force / polynomial fit



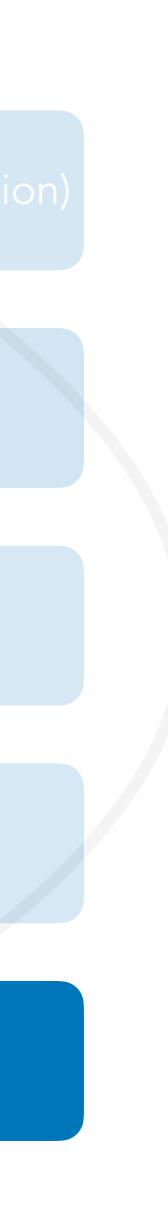
nput Current Data (Func

Reduce variables by dimensionality

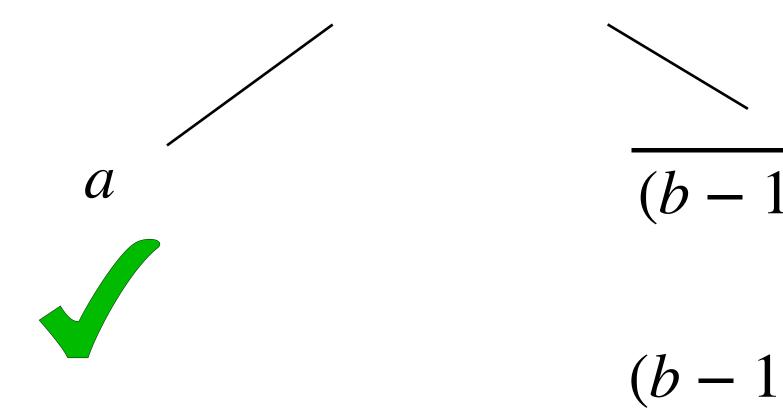
Train a NN

Query NN to reduce variables

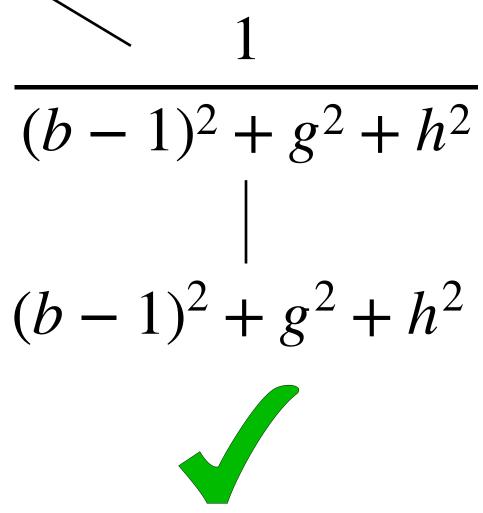
Poly fit / Brute force

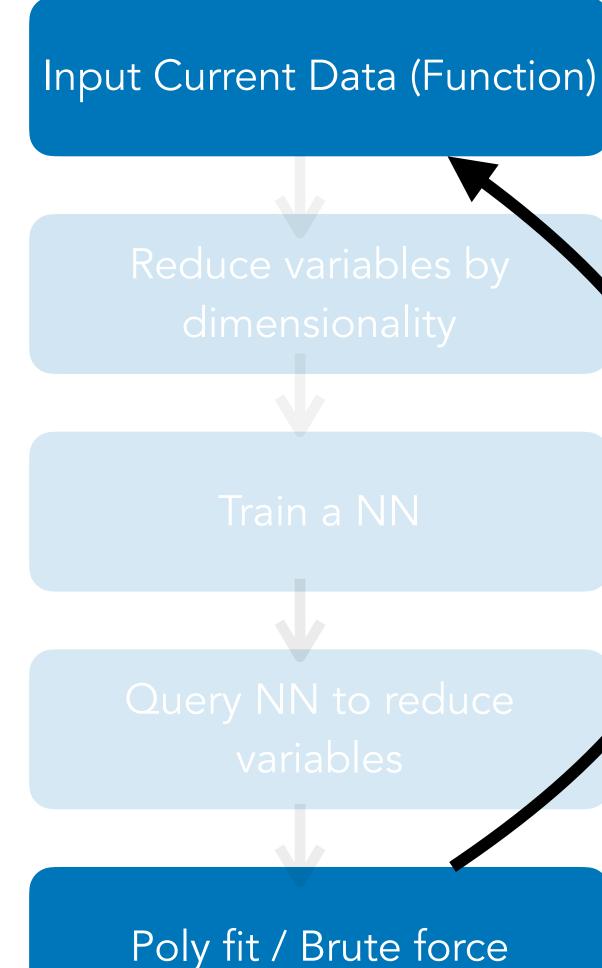


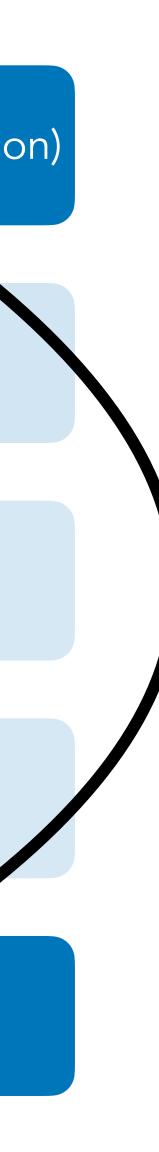
Test brute force / polynomial fit



If this fails, then recursively take the smaller problems and rerun







This was tested on 100 equations from Feynman's lecture notes and 20 harder equations

| Feynman | Equation | Feynman | Equation |
|---------|--|----------|--|
| c | | eq. | |
| eq. | $-\theta^2/2/\sqrt{2}$ | II.2.42 | $P = \frac{\kappa (T_2 - T_1)A}{P_E}$ $F_E = \frac{P^d}{4\pi r^2}$ $V_e = \frac{q}{4\pi \epsilon r}$ |
| I.6.20a | $\begin{vmatrix} f = e^{-\theta^2/2} / \sqrt{2\pi} \\ f = e^{-\frac{\theta^2}{2\sigma^2}} / \sqrt{2\pi\sigma^2} \end{vmatrix}$ | II.3.24 | $F_E = \frac{P^{-1}}{4\pi r^2}$ |
| I.6.20 | $\int f = e^{-\frac{\theta^2}{2\sigma^2}} / \sqrt{2\pi\sigma^2}$ | II.4.23 | $V_e = \frac{q}{4\pi\epsilon r}$ |
| 1.0.20 | | II.6.11 | $V_e = \frac{1}{4\pi\epsilon} \frac{p_d \cos \theta}{r^2}$ |
| I.6.20b | $\int f = e^{-\frac{(\theta - \theta_1)^2}{2\sigma^2}} / \sqrt{2\pi\sigma^2}$ | II.6.15a | $E_f = \frac{3}{4\pi\epsilon} \frac{p_d z}{r^5} \sqrt{x^2 + y^2}$ |
| I.8.14 | $d = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2}$ | II.6.15b | $E_f = \frac{\frac{3}{4\pi\epsilon}}{\frac{1}{4\pi\epsilon}} \frac{\frac{p_d z}{r^5}}{r^5} \sqrt{x^2 + y^2}$ $E_f = \frac{\frac{3}{4\pi\epsilon}}{\frac{1}{4\pi\epsilon}} \frac{\frac{p_d}{r^3}}{r^3} \cos\theta \sin\theta$ |
| I.9.18 | $F = \frac{Gm_1m_2}{(r_0 - r_1)^2 + (r_0 - r_1)^2 + (r_0 - r_1)^2}$ | II.8.7 | $E = \frac{3}{5} \frac{q^2}{4\pi\epsilon d}$ $E_{den} = \frac{\epsilon E_f^2}{2}$ |
| I.10.7 | $m = \frac{m_0^{(m_1)} + (g_2 - g_1) + (z_2 - z_1)}{\sqrt{m_0}}$ | II.8.31 | $E_{den} = \frac{\epsilon E_{f}}{2}$ |
| | $m = \frac{\frac{w_2 w_1}{m_0}}{\sqrt{1 - \frac{w^2}{c^2}}}$ | II.10.9 | $E_f = \frac{\sigma_{den}}{\epsilon} \frac{1}{1+\chi}$ |
| I.11.19 | $A = x_1 y_1 + x_2 y_2 + x_3 y_3$ | II.11.3 | $E_f = \frac{\frac{\sigma_{den}}{\epsilon}}{\frac{qE_f}{m(\omega_0^2 - \omega^2)}}$ |
| I.12.1 | $F = \mu N_n$ | II.11.17 | $n = n_0 (1 + \frac{p_d E_f \cos \theta}{k_b T})$ |
| I.12.2 | $F = \frac{q_1 q_2}{A \pi \epsilon r^2}$ | | $P_* = \frac{n_\rho p_d^2 E_f}{3k_b T}$ |
| I.12.4 | $\begin{vmatrix} F = \frac{q_1 q_2}{4\pi \epsilon r^2} \\ E_f = \frac{q_1}{4\pi \epsilon r^2} \end{vmatrix}$ | | 0 |
| I.12.5 | $F = q_2 E_f$ | II.11.27 | $P_* = \frac{n\alpha}{1 - n\alpha/3} \epsilon E_f$ |
| I.12.11 | $F = q(E_f + Bv\sin\theta)$ | 11.11.28 | $\theta = 1 + \frac{n\alpha}{1 - (n\alpha/3)}$ $B = \frac{1}{4\pi\epsilon c_{\alpha}^2} \frac{2I}{r}$ |
| I.13.4 | $K = \frac{1}{2}m(v^2 + u^2 + w^2)$ | II.13.17 | $B = \frac{1}{4\pi\epsilon c_{0,0}^2} \frac{21}{r}$ |
| I.13.12 | $U = Gm_1m_2(\frac{1}{r_2} - \frac{1}{r_1})$ | II.13.23 | $ ho_c = rac{ ho_c_0}{\sqrt{1 - v^2/c^2}}$ |
| I.14.3 | U = mgz | II.13.34 | $\rho_{c} = \frac{\frac{4\pi c c_{\rho_{c_{0}}}}{\sqrt{1 - v^{2}/c^{2}}}}{j = \frac{\rho_{c_{0}} v}{\sqrt{1 - v^{2}/c^{2}}}}$ |
| | | II.15.4 | $E = -\mu_M B \cos \theta$ |
| I.14.4 | $U = \frac{k_{spring}x^2}{2}$ | II.15.5 | $E = -p_d E_f \cos \theta$ |
| I.15.3x | $x_1 = \frac{x-u}{\sqrt{1-u^2/c^2}}$ | II.21.32 | $V_e = \frac{\frac{1}{q}}{\frac{q}{4\pi\epsilon r(1-v/c)}}$ |
| I.15.3t | $\begin{vmatrix} x_1 = \frac{x^2 - ut}{\sqrt{1 - u^2/c^2}} \\ t_1 = \frac{t - ux/c^2}{\sqrt{1 - u^2/c^2}} \end{vmatrix}$ | II.24.17 | $k = \sqrt{rac{\omega^2}{c^2} - rac{\pi^2}{d^2}}$ |
| I.15.10 | $\Big _{n} - \frac{\sqrt{1-u^2/c^2}}{m_0 v}$ | II.27.16 | $F_E = \epsilon c E_f^2$ |
| 1.15.10 | $p = \frac{1}{\sqrt{1 - v^2/c^2}}$ | II.27.18 | $E_{den} = \epsilon E_f^2$ |
| I.16.6 | $\begin{vmatrix} p = \frac{\sqrt{\frac{1-u}{m_0 v}}}{\sqrt{1-v^2/c^2}} \\ v_1 = \frac{u+v}{1+uv/c^2} \end{vmatrix}$ | II.34.2a | $I = \frac{qv}{2\pi r}$ |
| I.18.4 | $r = \frac{\frac{m_1 r_1 + m_2 r_2}{m_1 + m_2}}{\frac{m_1 r_1 + m_2 r_2}{m_1 + m_2}}$ | II.34.2 | $\mu_M = \frac{qvr}{2}$ |
| | $m_1 + m_2$ | II.34.11 | $\begin{array}{l} \mu_M = \frac{qvr}{2} \\ \omega = \frac{g_qB}{2m} \end{array}$ |
| | | | I 11 I. |

| $ \begin{array}{ll} \mbox{Rutherford Scattering} & A = \left(\frac{Z_1 Z_2 \alpha \hbar c}{4E \sin^2 \left(\frac{\theta}{2}\right)}\right)^2 \\ \mbox{Friedman Equation} & H = \sqrt{\frac{8\pi G}{3}} \rho - \frac{k_f c^2}{a_f^2} \\ \mbox{Compton Scattering} & U = \frac{E}{1 + \frac{E}{mc^2} (1 - \cos \theta)} \\ \mbox{Radiated gravitational wave power} & P = -\frac{32}{5} \frac{G^4}{c^5} \frac{(m_1 m_2)^2 (m_1 + m_2)}{r^5} \\ \mbox{Relativistic aberration} & \theta_1 = \arccos\left(\frac{\cos \theta_2 - \frac{v}{c}}{1 - \frac{v}{c} \cos \theta_2}\right) \\ \mbox{N-slit diffraction} & I = I_0 \left[\frac{\sin(\alpha/2) \sin(N\delta/2)}{\alpha/2 \sin(\delta/2)}\right]^2 \\ \mbox{Goldstein 3.16} & v = \sqrt{\frac{2}{m} (E - U - \frac{L^2}{2mr^2})} \\ \mbox{Goldstein 3.64 (ellipse)} & r = \frac{d(1 - \alpha^2)}{1 + \alpha \cos(\theta_1 - \theta_2)} \\ \mbox{Goldstein 3.74 (Kepler)} & t = \frac{2\pi d^{3/2}}{\sqrt{G(m_1 + m_2)}} \end{array} $ |
|--|
| $\begin{array}{ll} \text{Compton Scattering} & U = \frac{E}{1 + \frac{E}{mc^2}(1 - \cos \theta)} \\ \text{Radiated gravitational wave power} & P = -\frac{32}{5}\frac{G^4}{c^5}\frac{(m_1m_2)^2(m_1 + m_2)}{r^5} \\ \text{Relativistic aberration} & \theta_1 = \arccos\left(\frac{\cos \theta_2 - \frac{v}{c}}{1 - \frac{v}{c}\cos \theta_2}\right) \\ \text{N-slit diffraction} & I = I_0 \left[\frac{\sin(\alpha/2)}{\alpha/2}\frac{\sin(N\delta/2)}{\sin(\delta/2)}\right]^2 \\ \text{Goldstein 3.16} & v = \sqrt{\frac{2}{m}(E - U - \frac{L^2}{2mr^2})} \\ \text{Goldstein 3.64 (ellipse)} & k = \frac{mk_G}{L^2}(1 + \sqrt{1 + \frac{2EL^2}{mk_G^2}}\cos(\theta_1 - \theta_2)) \\ & r = \frac{d(1 - \alpha^2)}{1 + \alpha\cos(\theta_1 - \theta_2)} \end{array}$ |
| $\begin{array}{ll} \text{Compton Scattering} & U = \frac{E}{1 + \frac{E}{mc^2}(1 - \cos \theta)} \\ \text{Radiated gravitational wave power} & P = -\frac{32}{5}\frac{G^4}{c^5}\frac{(m_1m_2)^2(m_1 + m_2)}{r^5} \\ \text{Relativistic aberration} & \theta_1 = \arccos\left(\frac{\cos \theta_2 - \frac{v}{c}}{1 - \frac{v}{c}\cos \theta_2}\right) \\ \text{N-slit diffraction} & I = I_0 \left[\frac{\sin(\alpha/2)}{\alpha/2}\frac{\sin(N\delta/2)}{\sin(\delta/2)}\right]^2 \\ \text{Goldstein 3.16} & v = \sqrt{\frac{2}{m}(E - U - \frac{L^2}{2mr^2})} \\ \text{Goldstein 3.64 (ellipse)} & k = \frac{mk_G}{L^2}(1 + \sqrt{1 + \frac{2EL^2}{mk_G^2}}\cos(\theta_1 - \theta_2)) \\ & r = \frac{d(1 - \alpha^2)}{1 + \alpha\cos(\theta_1 - \theta_2)} \end{array}$ |
| Radiated gravitational wave power $P = -\frac{32}{5} \frac{G^4}{c^5} \frac{(m_1 m_2)^2 (m_1 + m_2)}{r^5}$ Relativistic aberration $\theta_1 = \arccos\left(\frac{\cos\theta_2 - \frac{v}{c}}{1 - \frac{v}{c}\cos\theta_2}\right)$ N-slit diffraction $I = I_0 \left[\frac{\sin(\alpha/2)}{\alpha/2} \frac{\sin(N\delta/2)}{\sin(\delta/2)}\right]^2$ Goldstein 3.16 $v = \sqrt{\frac{2}{m}(E - U - \frac{L^2}{2mr^2})}$ Goldstein 3.55 $k = \frac{mk_G}{L^2} (1 + \sqrt{1 + \frac{2EL^2}{mk_G^2}}\cos(\theta_1 - \theta_2))$ Goldstein 3.64 (ellipse) $r = \frac{d(1 - \alpha^2)}{1 + \alpha\cos(\theta_1 - \theta_2)}$ |
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| Goldstein 3.64 (ellipse) $r = \frac{a(1-\alpha^{-1})}{1+\alpha\cos(\theta_1-\theta_2)}$ |
| Goldstein 3.64 (ellipse) $r = \frac{a(1-\alpha^{-1})}{1+\alpha\cos(\theta_1-\theta_2)}$ |
| Goldstein 3.64 (ellipse) $r = \frac{a(1-\alpha^{-1})}{1+\alpha\cos(\theta_1-\theta_2)}$ |
| Goldstein 3 74 (Kepler) $t = \frac{2\pi d^{3/2}}{2\pi d^{3/2}}$ |
| $c = \sqrt{G(m_1 + m_2)}$ |
| Goldstein 3.99 $\alpha = \sqrt{1 + \frac{2\epsilon^2 E L^2}{m(Z - Z - z^2)^2}}$ |
| Goldstein 8.56 $E = \sqrt{(p - qA_{vec})^2 c^2 + m^2 c^4} + qV_e$ |
| Goldstein 12.80 $E = \frac{1}{2m} \left[p^2 + m^2 \omega^2 x^2 (1 + \alpha \frac{x}{y}) \right]$ |
| Goldstein 8.56 Goldstein 12.80 Jackson 2.11 $V = m(Z_1Z_2q^2)^2$ $E = \sqrt{(p - qA_{vec})^2 c^2 + m^2 c^4} + qV_e$ $E = \frac{1}{2m} [p^2 + m^2 \omega^2 x^2 (1 + \alpha \frac{x}{y})]$ $F = \frac{q}{4\pi\epsilon y^2} \left[4\pi\epsilon V_e d - \frac{qdy^3}{(y^2 - d^2)^2} \right]$ |

This was tested on 100 equations from Feynman's lecture notes and 20 harder equations

The success rate was better than existing state of the art:

| Equation Set | Eureqa (Benchmark) | AI Feynman (with DA) | AI Feynman (without DA) |
|-------------------------------|--------------------|----------------------|-------------------------|
| 100 Feynman Lecture Equations | 68% | 100% | 93% |
| 20 "Bonus" Equations | 15% | 90% | ? |

Couple of interesting notes:

1. Failure cases. For example, the Radiational Gravitational Waves equation:

$$P = -\frac{32}{5} \frac{G^4}{c^5} \frac{(m_1 m_2)^2 (m_1 + m_2)}{r^5} \qquad \longrightarrow \qquad y = -\frac{32a^2 (1+a)}{5b^5}$$

In reverse polish notation, this is the string

Which would take too long to solve. This is separable, but the 5th power in the denominator cause a wide dynamic range and so it was not detected

aaa > * * *bbbbb* * * * * /

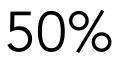


Couple of interesting notes:

with noise around $\epsilon = 10^{-2}$.

The thresholds were not adjusted to each problem so the tolerance can probably be improved.

2. Noise: Performance was consistent up to Gaussian noise levels of $\epsilon = 10^{-4}$ but dropped by 50%



Symbolic Regression - Example

Couple of interesting notes:

otherwise how we would discover relations such as $G\sqrt{\pi}$, $\frac{G^4}{c^5}$, etc.

But how do we learn these specific relationships with real data?

3. Constants: Constants were treated as their own variables and symbols in the data. It's not clear



Symbolic Regression

Problems:

- 1. Sensitivity to noise
- 2. Constant Relationships are difficult to detect
- 3. Require lots of data to train a neural network

Incorporate physics knowledge into the search

Approach 2: Symbolic Regression with background knowledge

Approach 2: Symbolic Regression with background knowledge

In particular, we will very briefly look at the approach in this paper:

nature communications

Article

https://doi.org/10.1038/s41467-023-37236-y

6

Combining data and theory for derivable scientific discovery with AI-Descartes

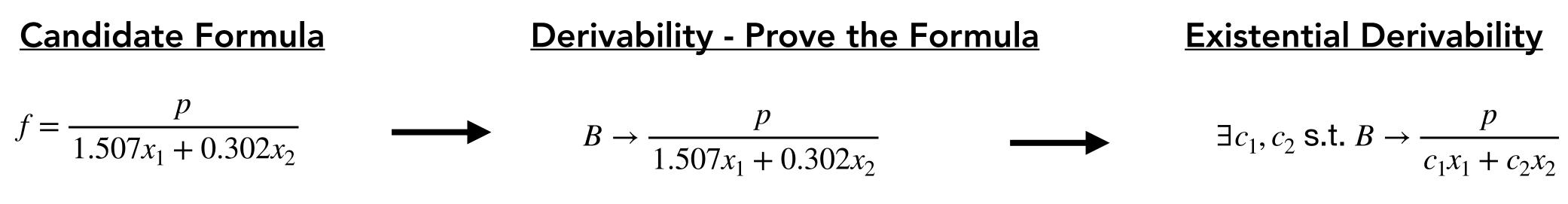
| Received: 28 October 2021 | Cristina Cornelio ^{1,2} , Sanjeeb Dash ¹ , Vernon Austel ¹ , Tyler R. Josephson ^{3,4} , Joao Goncalves ¹ , Kenneth L. Clarkson ¹ , Nimrod Megiddo ¹ , Bachir El Khadir ¹ & Lior Horesh ^{1,5} | | | |
|---------------------------------|---|--|--|--|
| Accepted: 8 March 2023 | | | | |
| Published online: 12 April 2023 | | | | |
| Check for updates | Scientists aim to discover meaningful formulae that accurately describe experimental data. Mathematical models of natural phenomena can be manually created from domain knowledge and fitted to data, or, in contrast, created automatically from large datasets with machine-learning algorithms. The problem of incorporating prior knowledge expressed as constraints on the functional form of a learned model has been studied before, while finding models that are consistent with prior knowledge expressed via general logical axioms is an open problem. We develop a method to enable principled deri- vations of models of natural phenomena from axiomatic knowledge and experimental data by combining logical reasoning with symbolic regression. We demonstrate these concepts for Kepler's third law of planetary motion, Einstein's relativistic time-dilation law, and Langmuir's theory of adsorption. We show we can discover governing laws from few data points when logical reasoning is used to distinguish between candidate formulae having similar error on the data. | | | |

Key Idea:

Use symbolic regression methods to generate lots of candidates for the formula Given some background theory*B*, use an Automated TheoremProver (ATP) to rank and provethese candidates. Select the

Main Contributions:

1. Incorporating a derivability module into the search.



Confirm Functional Form

Run your favorite method of fitting parameters

Main Contributions:

2. If not derivable, then have the theorem proved derive bounds on the error between candidate function and functions derivable from the background theory

<u>Point wise</u> reasoning error

$$\beta_2^r = \sqrt{\sum_{i=1}^m \frac{f(x_i) - f_{\mathscr{B}}(x_i)}{f_{\mathscr{B}}(x_i)}} \qquad \beta_{\mathscr{B}}(x_i)$$

<u>Generalized</u> <u>reasoning error</u>

$$\beta_{\infty}^{r} = \sqrt{\sum_{i=1}^{m} \frac{f(x_{i}) - f_{\mathscr{B}}(x_{i})}{f_{\mathscr{B}}(x_{i})}}$$

Results: Ran this on more problems from the Feynman notes as well as supplementary problems.

| Label | Formula | AI-Descartes | AI Feynman | PySR | BMS |
|-----------|--|---------------------|----------------|----------------|----------------|
| II.11.20 | $n_{\rho}p_d^2 E_f/(3k_bT)$ | \checkmark^1 | х | х | \checkmark^2 |
| II.11.27 | $\frac{nlpha}{1-nlpha/3} \varepsilon E_f$ | \checkmark^1 | Х | Х | Х |
| II.11.28 | $1 + \frac{n\alpha}{1 - n\alpha/3}$ | \checkmark^1 | Х | Х | Х |
| II.13.17 | $\frac{1}{4\pi\varepsilon c^2} \frac{2I}{r}$ | \checkmark^1 | Х | \checkmark^1 | Х |
| П.13.23 | $\frac{4\pi \epsilon c_0^2}{\sqrt{1-v^2/c^2}}$ | \checkmark^1 | Х | Х | Х |
| II.13.34 | $\frac{\sqrt{\frac{\rho_{c_0}v}{\rho_{c_0}v}}}{\sqrt{1-v^2/c^2}}$ | Х | Х | Х | Х |
| П.21.32 | $\frac{\sqrt{q}}{4\pi\varepsilon r(1-v/c)}$ | Х | Х | Х | Х |
| II.24.17 | $\sqrt{rac{\omega^2}{c^2}-rac{\pi^2}{d^2}}$ | Х | Х | Х | Х |
| II.27.16 | $\varepsilon c E_f^2$ | \checkmark | Х | \checkmark | \checkmark |
| II.27.18 | $\epsilon E_f^{2'}$ | \checkmark^1 | \checkmark^1 | \checkmark | \checkmark |
| II.34.2a | $qv/(2\pi r)$ | \checkmark^1 | \checkmark^1 | \checkmark^1 | \checkmark^1 |
| II.34.2 | qvr/2 | \checkmark^1 | \checkmark^1 | \checkmark^1 | \checkmark^1 |
| II.34.11 | gqB/(2m) | \checkmark | \checkmark^1 | \checkmark^1 | \checkmark^1 |
| II.34.29a | $qh/(4\pi m)$ | \checkmark^1 | \checkmark^1 | \checkmark^1 | \checkmark^1 |
| II.34.29b | $g\mu_M BJ_z/h$ | \checkmark^2 | Х | \checkmark^1 | \checkmark^1 |
| II.35.18 | $\frac{\frac{n_0}{e^{\frac{mom \times B}{kb \times T}} + e^{\frac{-mom \times B}{(kb \times T)}}}$ | Х | Х | Х | Х |
| II.36.38 | $\frac{\mu_m B}{k_b T} + \frac{\mu_m \alpha M}{\varepsilon c^2 k_b T}$ | Х | Х | Х | Х |
| П.37.1 | $\mu_M(1+\chi)B$ | \checkmark^1 | Х | \checkmark^1 | \checkmark^2 |
| П.38.3 | YAx/d | \checkmark^1 | \checkmark^2 | \checkmark | \checkmark^2 |
| II.38.14 | $\frac{Y'}{2(1+\sigma)}$ | \checkmark^1 | \checkmark^2 | Х | \checkmark^2 |
| | Accuracy | 86.67% | 80% | 73.33% | 80% |

Some Comments:

Failure Cases:

Most of the failure cases came from when the theorem prover could not prove derivability or good bounds on the error.

Approach 3: Polynomial Optimization

Approach 3: Polynomial Optimization

In particular, we will look at the approach in this paper:

nature communications

Article

nttps://doi.org/10.1038/s41467-024-50074-w

Evolving scientific discovery by unifying data and background knowledge with AI Hilbert

| Received: 23 September 2023 | Ryan Cory-Wright ¹ , Cristina Cornelio ² , Sanjeeb Dash ³ , Bachir El Khadir ³ & | | |
|--------------------------------|--|--|--|
| Accepted: 27 June 2024 | Lior Horesh ^{® 3} | | |
| Published online: 14 July 2024 | | | |
| Check for updates | The discovery of scientific formulae that parsimoniously explain natural phenomena and align with existing background theory is a key goal in science. Historically, scientists have derived natural laws by manipulating equations based on existing knowledge, forming new equations, and verifying them experimentally. However, this does not include experimental data within the discovery process, which may be inefficient. We propose a solution to this problem when all axioms and scientific laws are expressible as polynomials and argue our approach is widely applicable. We model notions of minimal complexity using binary variables and logical constraints, solve polynomial optimization problems via mixed-integer linear or semidefinite optimization, and prove the validity of our scientific discoveries in a principled manner using Positivstellensatz certificates. We demonstrate that some famous scientific laws, including Kepler's Law of Planetary Motion and the Radiated Gravitational Wave Power equation, can be derived in a principled manner from axioms and experimental data. | | |

9

Key Idea: There are a lot of laws in science that can be transformed to polynomial expressions.

Examples:

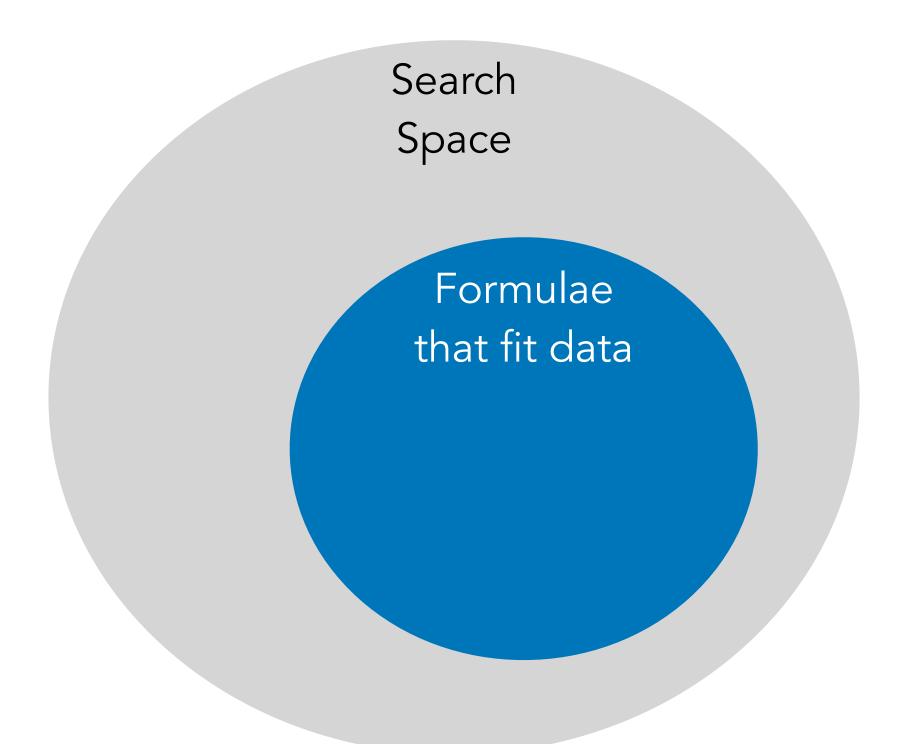
Kepler's Third Law:
$$p = \sqrt{\frac{4\pi^2(d_1+d_2)^3}{G(m_1+m_2)}} \mapsto p^2 G(m_1+m_2) - 4\pi^2(d_1+d_2)^3$$

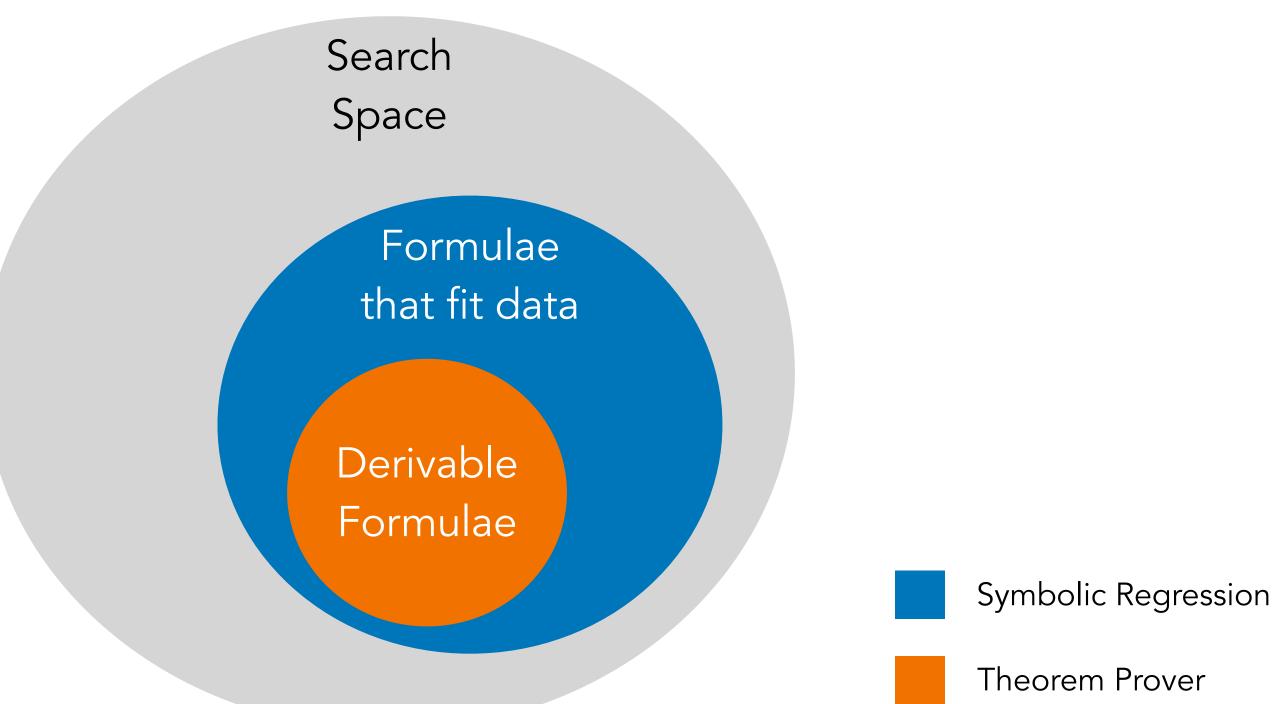
Compton Scattering Equation: $\lambda_2 - \lambda_1 = \frac{h}{m_e c}(1 - \cos\theta) \mapsto (\lambda_2 - \lambda_1)hm_e c^3(1 + \cos\theta)$

Note: We will model things after polynomials and polynomial inequalities. We will not deal with diffeqs and trig inequalities directly, but we can get some mileage as we will see.

Key Idea: There are a lot of laws in science that can be transformed to polynomial expressions.

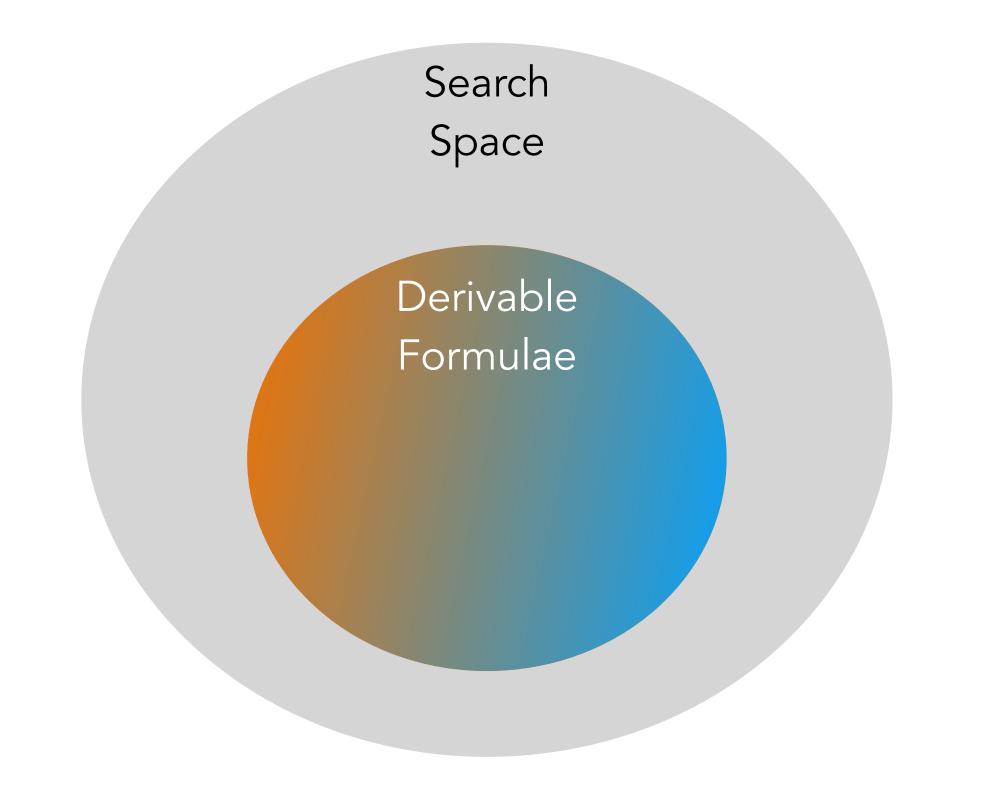
Previous methods:





Key Idea: There are a lot of laws in science that can be transformed to polynomial expressions.

This work:



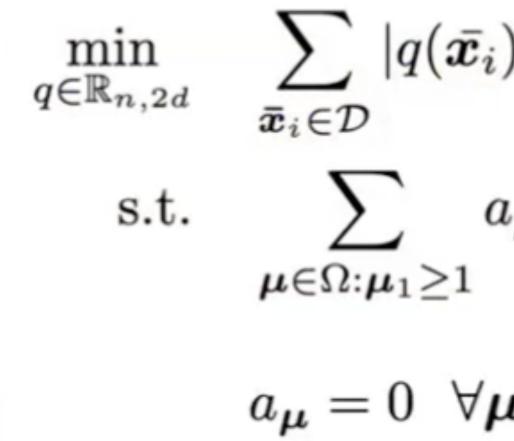


Polynomial Optimization



Algebraic Geometry Results



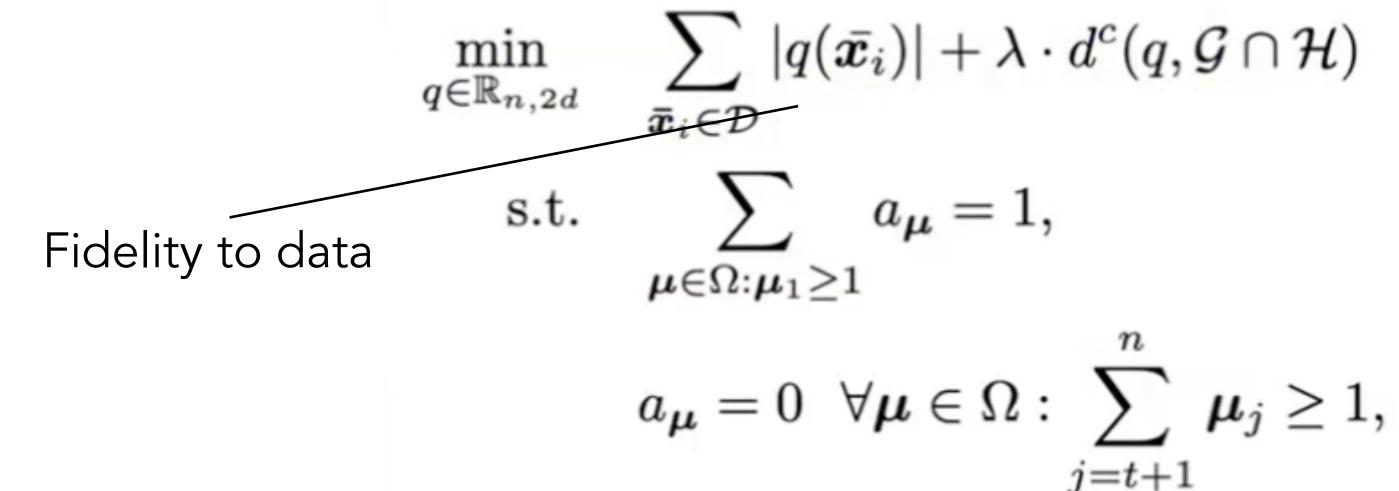


$$|| + \lambda \cdot d^{c}(q, \mathcal{G} \cap \mathcal{H})|$$

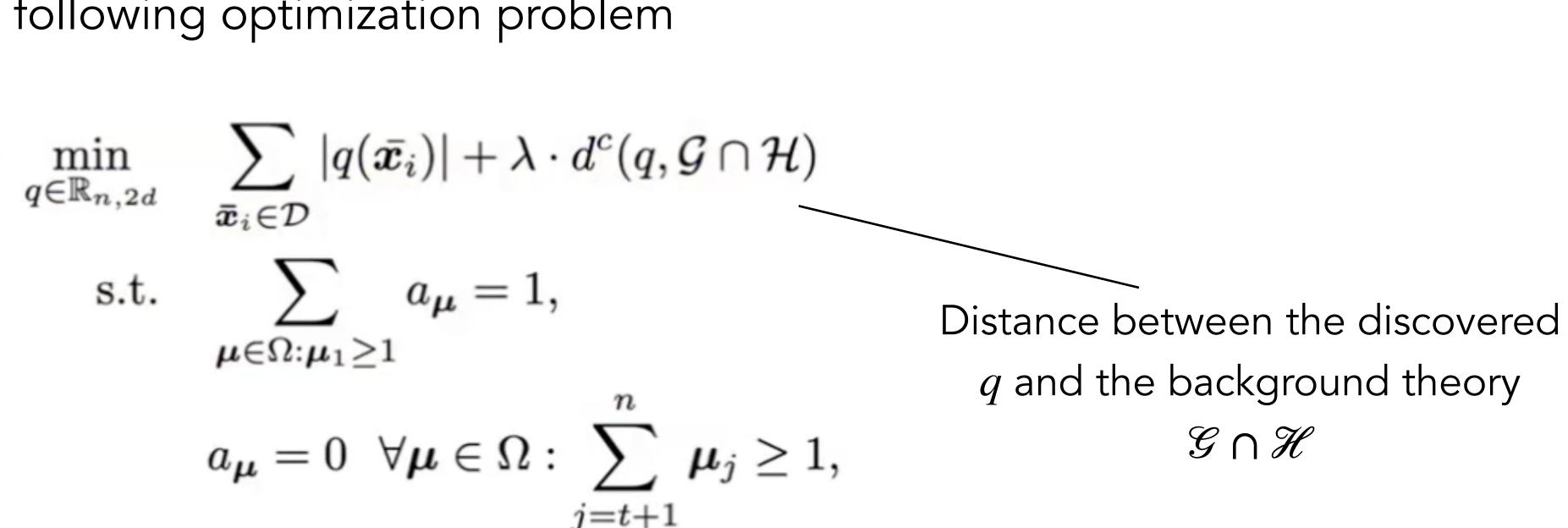
 $a_{\mu} = 1,$

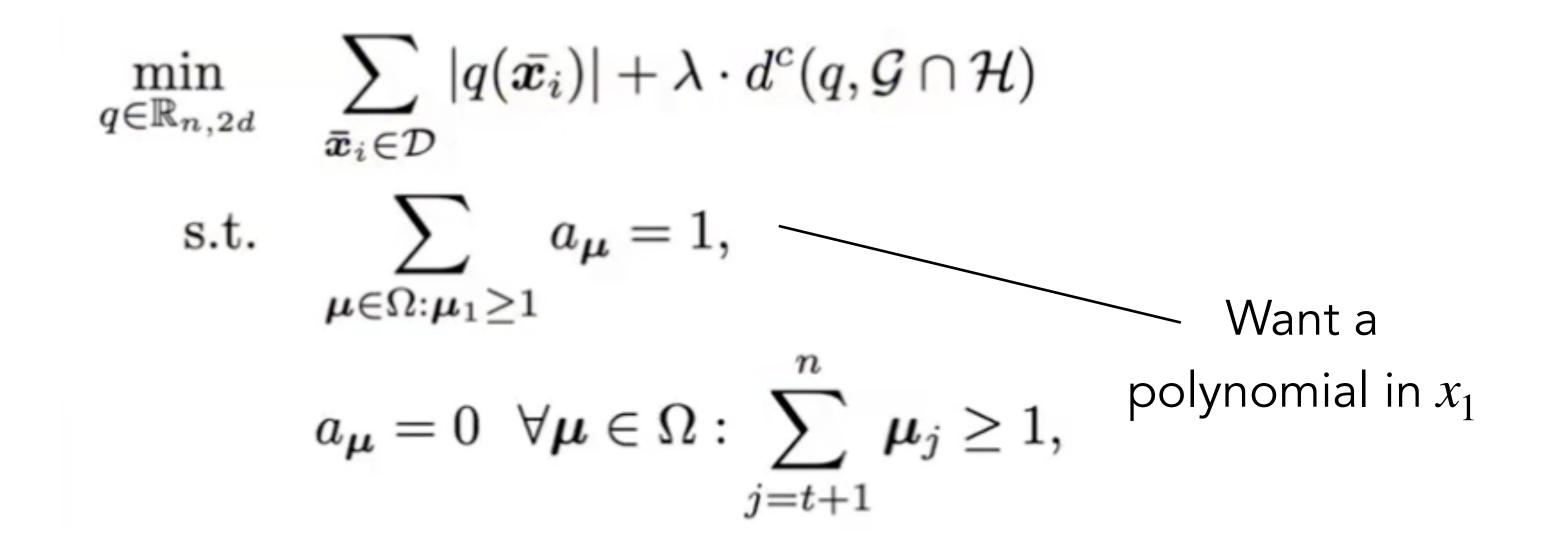
$$\boldsymbol{\mu} \in \Omega: \sum_{j=t+1}^{n} \boldsymbol{\mu}_j \ge 1,$$

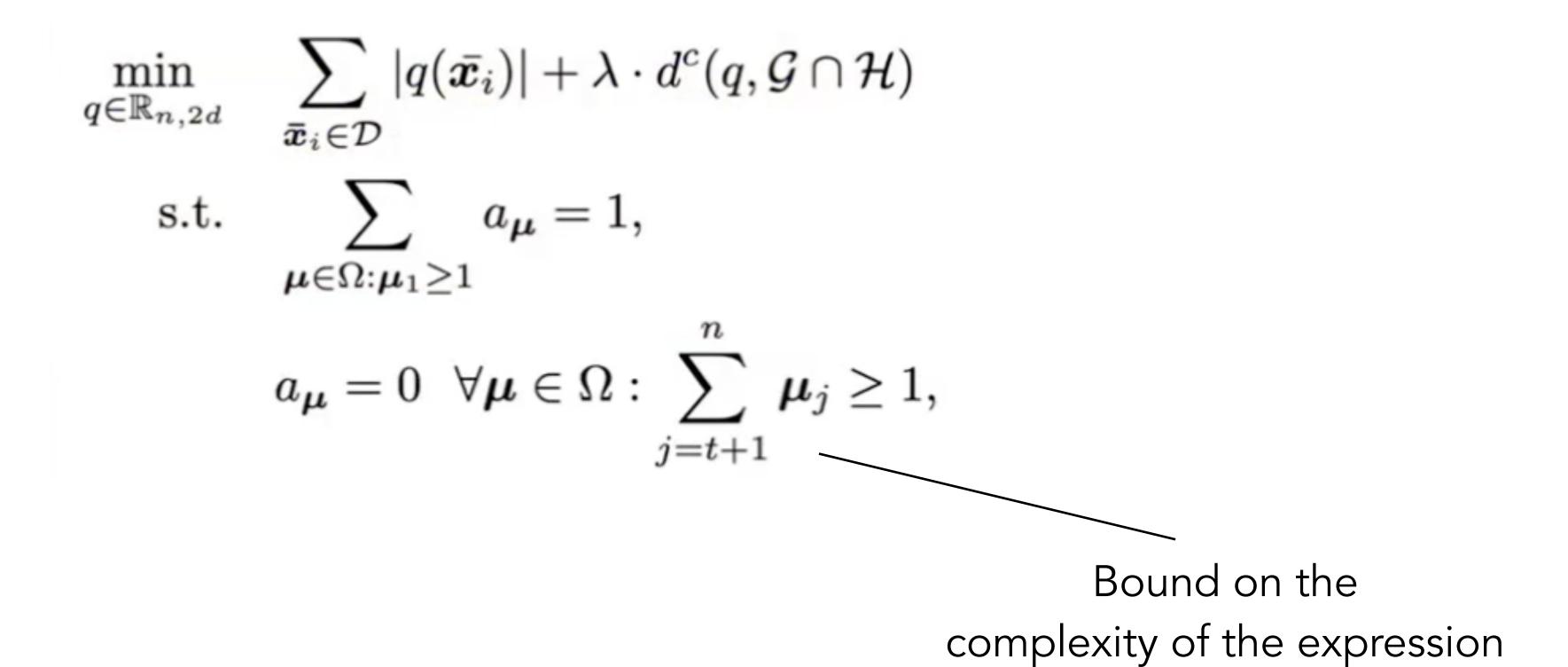
The idea is to solve the following optimization problem



This can be defined with l_2 or l_{∞} loss







Polynomial Optimization - Distance to Background

Given some axioms $h_1(\mathbf{x}), \ldots, h_l(\mathbf{x}) = 0$, inequalities $g_1(\mathbf{x}) \ge 0, \ldots, g_k(\mathbf{x}) \ge 0$, where $g_i, h_j \in \mathbb{R}[x_1, \dots, x_n]$ and data \mathcal{D} , define the sets

$$\mathscr{G} = \{ x \in \mathbb{R}^n \,|\, g_i(x) = 0 \}$$

We want to find a polynomial f such that

 $x \in \mathscr{G} \cap$

<u>Theorem: Putinar's Positive Stellensatz</u>

If g_i, h_k satisfy the Archimedean property, i.e. there exists R and sum of squares polynomials $\alpha_0, \ldots, \alpha_l$ such that squares polynomials $\alpha_0, \ldots, \alpha_k$ and real polynomials β_1, \ldots, β_l such that

 $f = \alpha_0$

0} and $\mathcal{H} = \{x \in \mathbb{R}^n | h_i(x) \ge 0\}$

$$\mathscr{H} \implies f(x) \ge 0$$

 $R - \sum x_i^2 = \alpha_0 + \sum \alpha_i g_i$, then for any degree d polynomial f, the implication above holds if and only if there exist sum of

$$+\sum_{i=1}^k \alpha_i g_i + \sum_{j=1}^l \beta_j h_j$$



Polynomial Optimization - Distance to Background

Given some axioms $h_1(\mathbf{x}), \ldots, h_l(\mathbf{x}) = 0$, inequalities $g_1(\mathbf{x}) \ge 0, \ldots, g_k(\mathbf{x}) \ge 0$, where $g_i, h_j \in \mathbb{R}[x_1, \dots, x_n]$ and data \mathcal{D} , define the sets

$$\mathcal{G} = \{ x \in \mathbb{R}^n \,|\, g_i(x) = 0 \}$$

We want to find a polynomial f such that

 $x \in \mathscr{G} \cap$

Case 1: Incomplete background knowledge

$$d(q, \mathcal{G} \cap \mathcal{H}) = \min_{\alpha_i \in \Sigma_{n,2d}, \beta_j \in \mathbb{R}_n, d} \|\text{Coeff} (q - \alpha_0 - \sum \alpha_i g_i - \sum \beta_j h_j)\|_2$$

0} and $\mathscr{H} = \{x \in \mathbb{R}^n | h_i(x) \ge 0\}$

$$\mathcal{H} \implies f(x) \ge 0$$

Polynomial Optimization - Distance to Background

Given some axioms $h_1(\mathbf{x}), \ldots, h_l(\mathbf{x}) = 0$, inequalities $g_1(\mathbf{x}) \ge 0, \ldots, g_k(\mathbf{x}) \ge 0$, where $g_i, h_i \in \mathbb{R}[x_1, \dots, x_n]$ and data \mathcal{D} , define the sets

$$\mathcal{G} = \{ x \in \mathbb{R}^n \,|\, g_i(x) = 0 \}$$

We want to find a polynomial f such that

 $x \in \mathscr{G} \cap$

Case 2: Inconsistent background knowledge

$$\begin{split} d(q, \mathscr{G} \cap \mathscr{H}) &= \min_{\alpha_i \in \Sigma_{n,2d}, \beta_j \in \mathbb{R}_n, d} \| \text{Coeff} \left(q - \alpha_0 - \sum \alpha_i g_i - \sum \beta_j h_j \right) \|_2 \\ \text{s.t. } \alpha_i &= 0 \text{ if } z_i = 0, z_i \in \{0, 1\} \\ \beta_j &= 0 \text{ if } y_j = 0, y_j \in \{0, 1\} \\ \sum z_i + \sum y_j &\leq \tau \end{split}$$

0} and $\mathcal{H} = \{x \in \mathbb{R}^n \mid h_i(x) \ge 0\}$

$$\mathcal{H} \implies f(x) \geq 0$$

Let's say we want to discover Kepler's third law

- $p = \sqrt{\frac{4(d_1 + d_2)^3}{G(m_1 + m_2)}}$ We have data $\{(d_1, d_2, m_1, m_2, p)_i\}$ as well as some knowledge of physics $(d_1 + d_2)^2 F_g - Gm_1 m_2 = 0,$ $F_c - m_2 d_2 w^2 = 0$,
- Then running solving the optimization problem on a solver gives
- Which is the correct expression post-factoring

 $d_1m_1 - d_2m_2 = 0$,

 $F_c - F_q = 0,$

wp = 1,

 $m_1 m_2 G p^2 - m_1 d_1 d_2^2 - m_2 d_1^2 d_2 - 2m_2 d_1 d_2^2 = 0$

Then running solving the optimization problem on a solver gives

 $m_1 m_2 G p^2 - m_1 d_1 d_1$

Which is the correct expression post-factoring. What's more is that we also get the corresponding α_i :

 $m_1d_1d_2^2pw + m_2d_1^2d_2pw + 2m_2d_1d_2^2p$

$$d_2^2 - m_2 d_1^2 d_2 - 2m_2 d_1 d_2^2 = 0$$

$$egin{aligned} & -d_2^2 p^2 w^2, \ & -p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & w + m_1 d_1 d_2^2 + m_2 d_1^2 d_2 + 2m_2 d_1 d_2^2, \end{aligned}$$

Then running solving the optimization problem on a solver gives

 $m_1 m_2 G p^2 - m_1 d_1 d_1$

Which is the correct expression post-factoring. What's more is that we also get the corresponding α_i :

 $m_1d_1d_2^2pw + m_2d_1^2d_2pw + 2m_2d_1d_2^2p$

$$d_2^2 - m_2 d_1^2 d_2 - 2m_2 d_1 d_2^2 = 0$$

$$egin{aligned} & -d_2^2 p^2 w^2, \ & -p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & d_1^2 p^2 + 2d_1 d_2 p^2 + d_2^2 p^2, \ & w + m_1 d_1 d_2^2 + m_2 d_1^2 d_2 + 2m_2 d_1 d_2^2, \end{aligned}$$

Some other examples that were tested:

Radiational Gravita

ational Wave Power:
$$P = -\frac{32G^4}{5c^5r^5}(m_1m_2)^2(m_1+m_2)$$
 with background
 $\omega^2 r^3 - G(m_1+m_2) = 0,$
 $5(m_1+m_2)^2 c^5 P + G \operatorname{tr} \left(\frac{d^3}{dt^3} \left(m_1m_2r^2 \begin{pmatrix} x^2 - \frac{1}{3} & xy & 0\\ xy & y^2 - \frac{1}{3} & 0\\ 0 & 0 & -\frac{1}{3} \end{pmatrix} \right)^2 \right) = 0,$
 $x^2 + y^2 = 1,$

Hagen-Poiseulle Equation: $u(r) = \frac{-\Delta p}{4L\mu}(r^2 - R^2)$

u = $\mu \frac{\partial}{\partial r} (r \frac{\partial}{\partial r} u)$

 $c_0 + L \frac{a}{c_0}$

$$= c_0 + c_2 r^2,$$

$$- r \frac{dp}{dx} = 0,$$

$$+ c_2 R^2 = 0,$$

$$\frac{dp}{dx} = -\Delta p,$$

Polynomial Optimization - Results

| Data | | AI-Hilbert | AI-Descartes [13] | AI-Feynman [42] | PySR [15] | BMS [24] |
|--------------|------------------------------------|--------------|-------------------|-----------------|--------------|----------|
| S | Hagen Poiseuille | \checkmark | × | × | × | × |
| \mathbf{S} | Gravitational Wave Power | \checkmark | × | × | × | × |
| R | Relat. Time Dilat. | \checkmark | × | × | × | × |
| R | Kepler's 3 Law | \checkmark | ✓* | × | \checkmark | × |
| \mathbf{S} | Bell inequalities [†] | \checkmark | × | × | × | × |
| S | [3.1] I.15.10 FSRD | \checkmark | √* | × | × | × |
| \mathbf{S} | [3.1] I.27.6 FSRD | \checkmark | ✓* | √* | \checkmark | × |
| \mathbf{S} | [3.1] I.34.8 FSRD | \checkmark | \checkmark | √* | \checkmark | √* |
| \mathbf{S} | [3.1] I.43.16 FSRD | \checkmark | ✓* | √* | \checkmark | √* |
| \mathbf{S} | [3.1] II.10.9 FSRD | \checkmark | ✓* | √* | √* | √* |
| \mathbf{S} | [3.1] II.34.2 FSRD | \checkmark | ✓* | √* | √* | √* |
| S | [3.2] Inelastic Relativ. Collision | ~ | ✓* | × | × | × |
| \mathbf{S} | [3.3] Decay of Pion ^{\$} | \checkmark | ✓* | × | × | √* |
| \mathbf{S} | [3.4] Radiation Damping | \checkmark | × | × | √* | √* |
| \mathbf{S} | [3.5] Escape Velocity | \checkmark | \checkmark | ✓* | √* | √* |
| \mathbf{S} | [3.6] Hall Effect | \checkmark | ✓* | × | √* | √* |
| S | [3.7] Compton Scattering | \checkmark | ✓* | ✓* | √* | √* |

Notation used in table:

✓ denotes the successful recovery of a scientific law.
 ✗ denotes the failure to recover a scientific law.
 ✓* denotes recovery up to constants, but not exact recovery.
 ✓[▶] denotes successful recovery when some variables that are potentially not observable are still declared to be observable.

Some thoughts:

1. Solving the LP can be slow. In the worst case (for gravitational waves)

Ongoing work

Projecting varieties

Aim: Discover Kepler's Third Law using just background theory

$$d_{1}m_{1} - d_{2}m_{2} = 0,$$

$$(d_{1} + d_{2})^{2}F_{g} - Gm_{1}m_{2} = 0,$$

$$F_{c} - m_{2}d_{2}w^{2} = 0,$$

$$F_{c} - F_{g} = 0,$$

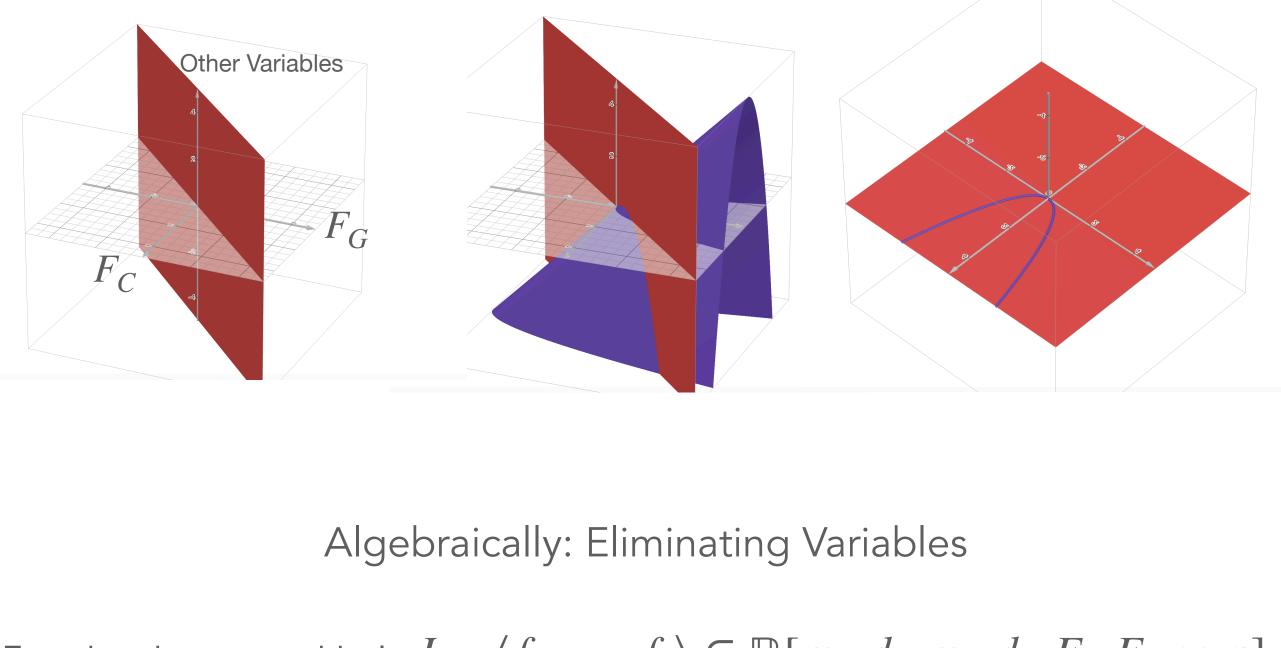
$$wp = 1,$$

$$p = \sqrt{\frac{4\pi^{2}(d_{1} + d_{2})^{3}}{G(m_{1} + m_{2})}},$$

$$mp = 1,$$

$$\pi : V(f_{1}, f_{2}, f_{3}, f_{4}, f_{5}) \subseteq \mathbb{R}^{8} \longrightarrow \bar{V} \subseteq \mathbb{R}^{5}$$

This can be done using Groebner bases:



Encode axioms as an ideal: $I = \langle f_1, \dots, f_5 \rangle \subseteq \mathbb{R}[m_1, d_1, m_2, d_2, F_c, F_g, w, p]$ Compute the intersection $I \cap \mathbb{R}[m_1, d_1, m_2, d_2, p]$

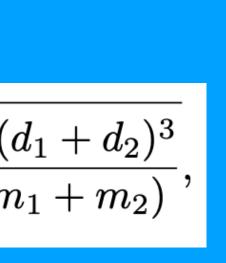
Encode axioms as an ideal: $G = \langle g_1, \ldots, g_k \rangle = I$

Compute the intersection $G \cap \mathbb{R}[m_1, d_1, m_2, d_2, p]$

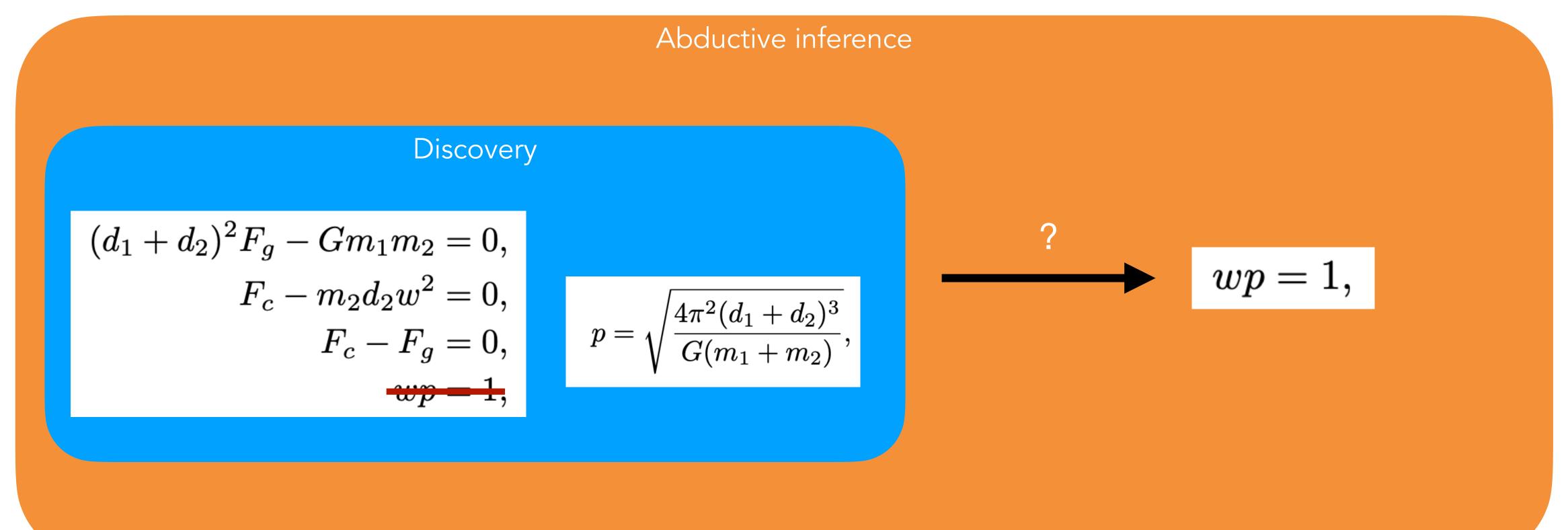
We know we can recover a formula from a complete axiom system

Discovery
$$(d_1+d_2)^2 F_g - Gm_1m_2 = 0,$$

 $F_c - m_2d_2w^2 = 0,$
 $F_c - F_g = 0,$
 $wp = 1,$
 $p = \sqrt{\frac{4\pi^2(Gm_1)^2}{G(m_1)^2}}$

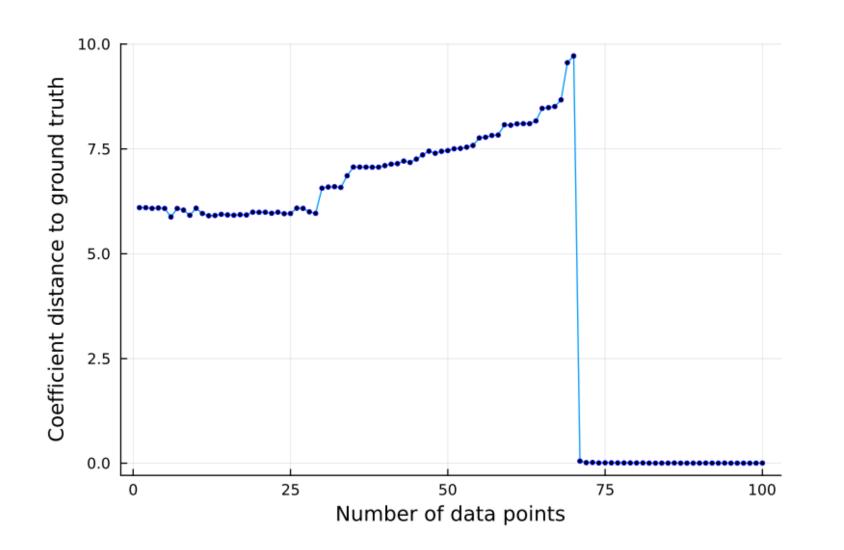


In practice, axiom systems can be inco axiom in the discovery process?



In practice, axiom systems can be incomplete. Can we rediscover a missing

Key idea: Using data and variable elimination, we can still discover formulae with AI Hilbert



Where $\alpha_4 F_4$ is unknown. Then $\langle f_1, f_2, f_3, q \rangle = \langle f_1, f_2, f_3, \alpha_4 F_4 \rangle$.



Rediscover Kepler with 70 data points if we omit wp - 1

$$p = \sqrt{\frac{4\pi^2(d_1 + d_2)^3}{G(m_1 + m_2)}},$$

Assumption: The discovered formula q can be proven if we had an additional axiom.

i.e.
$$q = \alpha_1 f_1 + \alpha_2 f_2 + \alpha_3 f_3 + \alpha_4 F_4$$

Take the new "discovered" surface and break it into its irreducible components

 $V(f_1, f_2, f_3, q) = \bigcup V(p_i)$

Where each $V(p_i)$ is irreducible. This is known as "primary" decomposition" in commutative algebra.

For Kepler's third law with wp - 1 missing, we get the following primary decomposition:

 $\langle f_1, f_2, f_3, q \rangle =$

 $\langle m_2, F_g, F_c \rangle \cap \langle d_2, m_1 \rangle$ $\langle F_g - Fc, (wp - 1), m_1 p^2 \rangle$ $\cap \langle F_g - F_c, (wp + 1), m_1 p \rangle$

$$\begin{array}{ccc} f_1 & (d_1+d_2)^2 F_g - Gm_1m_2 = 0, \\ f_2 & F_c - m_2 d_2 w^2 = 0, \\ f_3 & F_c - F_g = 0, \\ q & p^2 m_1 m_2 - d_1^2 d_2 m_2 - 2 d_1 d_2^2 m_2 - d_2^3 m_2 \end{array}$$

$$\begin{array}{l} _{1},F_{g},F_{c}\rangle \cap \left\langle d_{1}+d_{2},m_{1},F_{c}-F_{g},w^{2}m_{2}d_{2}-F_{g}\right\rangle \cap \\ -d_{1}^{2}d_{2}-2d_{1}d_{2}^{2}-d_{2}^{3},F_{g}(d_{1}+d_{2})^{2}-m_{1}m_{2},wm_{2}d_{2}-F_{g}p\rangle \cap \\ ^{2}-d_{1}^{2}d_{2}-2d_{1}d_{2}^{2}-d_{2}^{3},F_{g}(d_{1}+d_{2})^{2}-m_{1}m_{2},wm_{2}d_{2}+F_{g}p\rangle \end{array}$$

Abductive inference results

| Problem & Axiom $\#$ | Axiom | Recovered* |
|----------------------|----------------------------------|----------------|
| Kepler 1 | $(d_1 + d_2)^2 F_g - m_1 m_2$ | \checkmark |
| Kepler 2 | $F_c - m_2 d_2 w^2$ | \checkmark |
| Kepler 3 | $F_c - F_g$ | \checkmark |
| Kepler 4 | wp-1 | \checkmark |
| Compton 1 | $E_1 + Ee_1 - E_2 - Ee_2$ | \checkmark |
| Compton 2 | $E_1 - hf_1$ | \checkmark |
| Compton 3 | $E_2 - hf_2$ | \checkmark |
| Compton 4 | p_1c-hf_1 | \checkmark |
| Compton 5 | $p_2c - hf_2$ | \checkmark |
| Compton 6 | $\lambda_2 f_1 - c$ | \checkmark |
| Compton 7 | $\lambda_2 f_2 - c$ | \checkmark |
| Compton 8 | $Ee1-m_c^2$ | \checkmark |
| Compton 9 | $Ee2^2 - c^2 pe2^2 - me^2 c^4$ | \mathbf{X}^+ |
| Compton 10 | $pe2^2 - p2^2 - p1^2 + 2p1p2cos$ | \mathbf{X}^+ |
| Einstein 1 | $cdt_0 - 2*d$ | \checkmark |
| Einstein 2 | $4L^2 - 4d^2 - v^2 dt^2$ | \checkmark |
| Einstein 3 | $f_0 dt_0 - 1$ | \checkmark |
| Einstein 4 | fdt-1 | \checkmark |
| Einstein 5 | c*dt-2L | \checkmark |
| Escape Velocity 1 | $K_i - rac{1}{2}mv_e^2$ | \checkmark |
| Escape Velocity 2 | $\tilde{K_f} = 0$ | \checkmark |
| Escape Velocity 3 | $U_i r + GMm$ | \checkmark |
| Escape Velocity 4 | $U_f = 0$ | \checkmark |
| Escape Velocity 5 | $K_i + U_i - (K_f + U_f)$ | \checkmark |

Light Damping 1 Light Damping 2 Light Damping 3 Light Damping 4 Light Damping 5 Hagen Poiseuille 1 Hagen Poiseuille 2 Hagen Poiseuille 3 Hagen Poiseuille 4 Neutrino Decay 1 Neutrino Decay 2 Neutrino Decay 3 Neutrino Decay 4 Neutrino Decay 5 Hall Effect 1 Hall Effect 2 Hall Effect 3 Hall Effect 4 Hall Effect 5 Hall Effect 6 Hall Effect 7 Hall Effect 8 Hall Effect 9

Other goings on

Dataset and Benchmarking

Aim: Creating a synthetic dataset of polynomial axioms, data, and discovered formulae for benchmarking progress in the field

Aim: Apply the tools developed to research problems in physics to discover new scientific facts. Current work ongoing in cosmology.

New Discoveries

Non-polynomial Systems

Aim: Extending the current framework to systems involving differential equations, black box axioms, and trig functions



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